GAIA SCIENCE ALERTS

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7th Gaia Science Alerts Workshop, Utrecht, 7.Dec.2016

TRANSIENTS ZOO



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TRANSIENTS APPLICATION

Stellar astrophysics (SNe, CVs, MFlares, Be stars, RCrB, FUOri) Final stages of stellar evolution (CCSNe, CVs, SLSNe) **Rare stages of the stellar evolution (RCrB, FUOri) Distance indicators at local and cosmological scales (Novae, SNIa, SLSNe?) Stellar-mass black holes** (Microlensing, CCSNe) **Supermassive black holes in galaxies (TDEs, AGN flares)** Mass distribution in the Galaxy, extra-solar planets (Microlensing)

GAIA AND OTHER SURVEYS

	Gaia	OGLE-IV (Magellanic System only)	Catalina Sky Survey	PTF	LSST (from 2020??)								
deg² day-1	≈ 1230	150	1200	1000	5000								
Avg Cadence	≈ 30 days	5 days	14 days	5 days	4 days								
Limiting mag	~20.5(19)	21	19.5	21	r=24.7								
f sky	all sky	0.02	0.6	0.2	<0.48								
spectra astrometry													

GAIA SPECTRA



On the CCDs

- two low resolution spectrographs, R~100

Measurements - 1 dimensional for stars G>13 mag

CLASSIFICATION OF ALERTS Spectral classification - the "secret" power of Gaia







transients typing – low false-positives rate!

Gaia is equipped with low-resolution (R~100) spectrographs. They will allow for classifying transients into types and for supernovae they will provide estimates for redshift and epoch based on just a single observation!

Blagorodnova et al. 2014

CLASSIFICATION OF ALERTS Spectral classification - the "secret" power of Gaia



Blagorodnova et al. 2014

EXAMPLES ALERTS



raw Gaia data!

GAIA'S ASTROMETRY

Current Gaia's astrometric precision: 50 mas down to 20 mag



GAIA'S ASTROMETRY but possible ~1mas!

Nuclear transients classification



GAIA'S ASTROMETRY Astrometric microlensing recognising black hole lenses









see Kris Rybicki's talk

run 459 tstart 2015-04-05 08:25:31.587000 tend 2015-04-05 16:41:20.201000 1619 HPs completed for run 459



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filter search-leda (LW also runs three more other filters, with different purposes) brightness: 18.9 - 13.1 NewSource alerts only (of 84,774,199 total, of 2,593,333 alerting) 850444

cut1: galactic plane removal |b|>15 deg27276cut2: ecliptic plane removal |beta|>15 deg18956cut3: saturated SDSS star (flag+g<17) found nearby (30")</td>cut4: saturated GSC2 star (vpg<16 or bmag<16 or rmag<16) (30")</td>cut5: LEDA catalogue of galaxies has a galaxy nearby (30")234

234 left to eyeball

17254

13802



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further detailed inspection of candidates



checking Gaia BPRP spectra



checking other Gaia detections nearby



NOT and LT spectroscopic classification of supernovae Gaia15acz and Gaia15aek

ATel #7378; S. Mattila, J. Harmanen, T. Kangas (University of Turku), A. S. Piascik, C. Davis, I. A. Steele (Liverpool John Moores University), N. Blagorodnova, M. Fraser, H. Campbell, S. Hodgkin, N. Walton (University of Cambridge), L. Wyrzykowski (Warsaw University Observatory), E. Kankare, R. Kotak (Queen's University Belfast) on 13 Apr 2015; 11:49 UT

Supernova type IIP 2 weeks past max

=> Gaia15aek

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further detailed inspection of candidates



checking other Gaia detections nearby

checking Gaia BPRP spectra

SourceId= 6651971285633867648 hp5= 11816

mainfid

hp12 radec : 272.674339 -56.267402)

reftime(JD

24560001

ALIBRATED PHOTOMETR

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17.4

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17,65

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JD time -

3455000





false alert

-> spectrum suggests contamination from the host -> cross-matching problem

-> old source observed again with new sourceid

PUBLISHING

http://gsaweb.ast.cam.ac.uk/alerts/alert/Gaia16aye/

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PUBLISHING

http://gsaweb.ast.cam.ac.uk/alerts/alert/Gaia16aye/followup

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Observation date (TCB)

STATS 2015

1980 objects were discovered by PS1 (prof) 267 objects were discovered by CRTS (prof) 240 objects were discovered by OGLE (prof) 182 objects were discovered by ASAS-SN (prof) 135 objects were discovered by Dark Energy Survey (prof) 127 objects were discovered by PTF (prof) 112 objects were discovered by Gaia (prof) 87 objects were discovered by High Cadence Transient Survey (HiTS) (prof) 62 objects were discovered by MASTER (prof) 60 objects were discovered by Subaru/Hyper Suprime-Cam (prof) 49 objects were discovered by La Silla-QUEST (prof) 26 objects were discovered by Kamil Hornoch et al. 22 objects were discovered by THU-NAOC Transient Survey (TNTS) (prof) 20 objects were discovered by Italian Supernovae Search Project 19 objects were discovered by SPitzer InfraRed Intensive Transients Survey (prof) 17 objects were discovered by Tim Puckett, Jack Newton, et al. 16 objects were discovered by Koichi Itagaki 13 objects were discovered by LOSS (prof) 13 objects were discovered by Stu Parker 9 objects were discovered by Xing Gao et al. (prof) 5 objects were discovered by Berto Monard 5 objects were discovered by Ron Arbour 5 objects were discovered by SkyMapper (prof)

STATS 2016

3797 objects were discovered by PS1 (prof) 897 objects were discovered by Gaia (prof) 323 objects were discovered by CRTS (prof) 177 objects were discovered by ATLAS (prof) 177 objects were discovered by OGLE (prof) 155 objects were discovered by ASAS-SN (prof) 98 objects were discovered by PTF (prof) 62 objects were discovered by Dark Energy Survey (prof) 45 objects were discovered by MASTER (prof) 44 objects were discovered by PMO-Tsinghua Supernova Survey (PTSS) (prof) 35 objects were discovered by La Silla-QUEST (prof) 23 objects were discovered by TNTS (prof) 23 objects were discovered by Tim Puckett, Jack Newton, et al. 20 objects were discovered by Kamil Hornoch et al. 20 objects were discovered by Stu Parker 17 objects were discovered by LOSS (prof) 13 objects were discovered by Italian Supernovae Search Project 11 objects were discovered by Xing Gao et al. (prof) 9 objects were discovered by SkyMapper (prof) 7 objects were discovered by Koichi Itagaki 7 objects were discovered by SPitzer InfraRed Intensive Transients Survey (prof) 6 objects were discovered by Emmanuel Conseil 5 objects were discovered by Koichi Nishiyama and Fujio Kabashima 4 objects were discovered by Ron Arbour 3 objects were discovered by Leonardo Tartaglia 2 objects were discovered by GSST / Constantine Emmanouilidi / Emelie Selander 2 objects were discovered by Grzegorz Duszanowicz and Michal Zolnowski 2 objects were discovered by HST (prof) 0 shipsha see alisessee alise Tourshand and





Search by name or coordinates

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https://wis-tns.weizmann.ac.il/

TNS

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TNS

SN 2016ifk



OPTICON FOLLOW-UP NETWORK ~20 active partners, ~30000 data points collected 2014-2016



FOLLOW-UP CALIBRATION SERVER

gsaweb.ast.cam.ac.uk/followup/





MJD

GAIA16AYE (AYERS ROCK) First binary microlensing event outside of the Galactic Bulge Follow-up was essential!

IMAGE OF THE WEEK

FOLLOW-UP OPPORTUNITY OF A RARE MICROLENSING EVENT



Light curve of the microlensing event Gaia16aye, composed by data from Gaia (dark spots) and supported by data from ground-based follow-up telescopes (each colour indicates a different observatory). The solid black line shows the current best microlensing model computed by Przemek Mróz. The sharp rises are called caustic crossings as explained in the text below.

http://www.cosmos.esa.int/web/gaia/IoW 20161027



Date [2016]

Overview **Payload Module** Light curve of binary microlensing event detected by Gala. Credit: ESA/Gala/DPAC, P. Mroz, L. Service Module Wyrrykowski, K.A. Rybicki (Warsaw)

http://sci.esa.int/gaia/58546-gaia-spies-two-temporarily-magnified-stars/

Involves AAVSO (US), amateurs from Italy, schools via LCOGT/Cardiff. Will add German amateurs associations.

will be published as a scientific paper soon

NOTE ON ATELS

The Astronomer's Telegram

Gaia16bhp, Gaia16bhq, Gaia16bht and Gaia16bhj transients confirmed by Mercator/Maia imaging

ATel #9525; S. Blanco-Cuaresma, T. Semaan, M. Roelens, L. Palaversa, N. Mowlavi, L. Eyer (Department of Astronomy, University of Geneva, Switzerland) on 21 Sep 2016; 17:16 UT

The photometric calibrations were obtained using the Cambridge Photometric Calibration Server (CPCS), designed and maintained by Sergey Koposov and Lukasz Wyrzykowski. We acknowledge ESA Gaia, DPAC and the Photometric Science Alerts Team (http://gsaweb.ast.cam.ac.uk/alerts).

NEW CHALLENGES FOR 2017+

- lower thresholds more alerts
- need for full automatisation (downside: more false alerts)
- calibrated BP-RP: full exploitation of transients spectra for classification
- improved astrometry of transients, to ~1mas
- photometric follow-up towards robotisation
- spectroscopic follow-up expand network