

# Gaia Alert Spectroscopy with SPRAT at the Liverpool Telescope

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(Apologies for our absence!)  
(and thanks to our presenter...)

<http://telescope.livjm.ac.uk/>

# Liverpool Telescope

A 2.0 metre unmanned fully robotic telescope at the Observatorio del Roque de Los Muchachos on the Canary island of La Palma, Spain.

Owned and operated by Liverpool JMU, with financial support from STFC.



Optical Imaging: IO:O (10 x 10 arcmin)

IR Imaging: IO:I (6 x 6 arcmin)

Rapid Photometry: RISE (10 x 10 arcmin)

Polarimetry: RINGO3 (4 x 4 arcmin)

Medium Resolution Spectroscopy: FRODOSpec, R=2500, 5000 (380-900nm)

Low Resolution Optical Spectroscopy: SPRAT, R=350 (400-800nm)

Low Resolution UV-Optical Spectroscopy: LOTUS, R=350 (320-640nm)

# SPRAT – Spectrograph for the Rapid Acquisition of Transients

- Fast acquisition using own science camera
- High throughput (VPH grating)

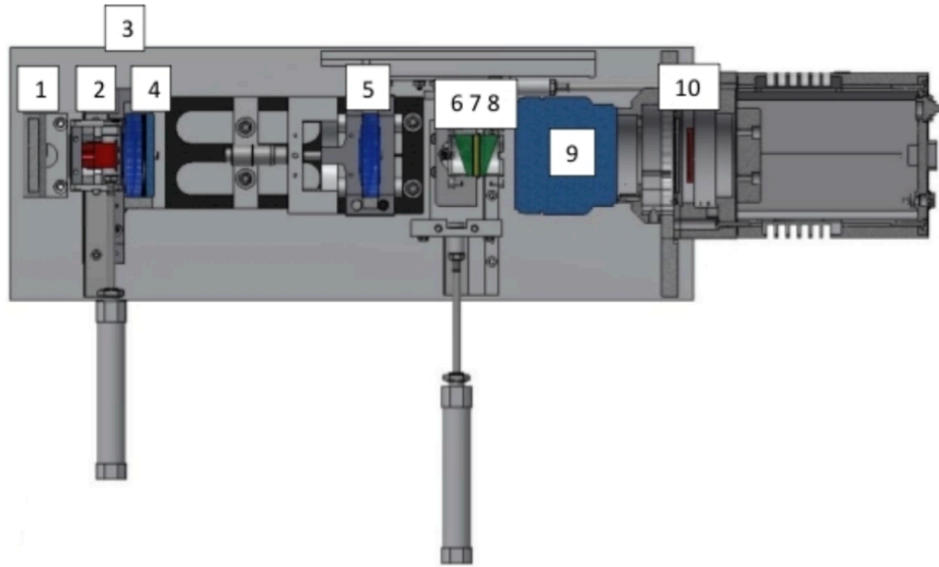
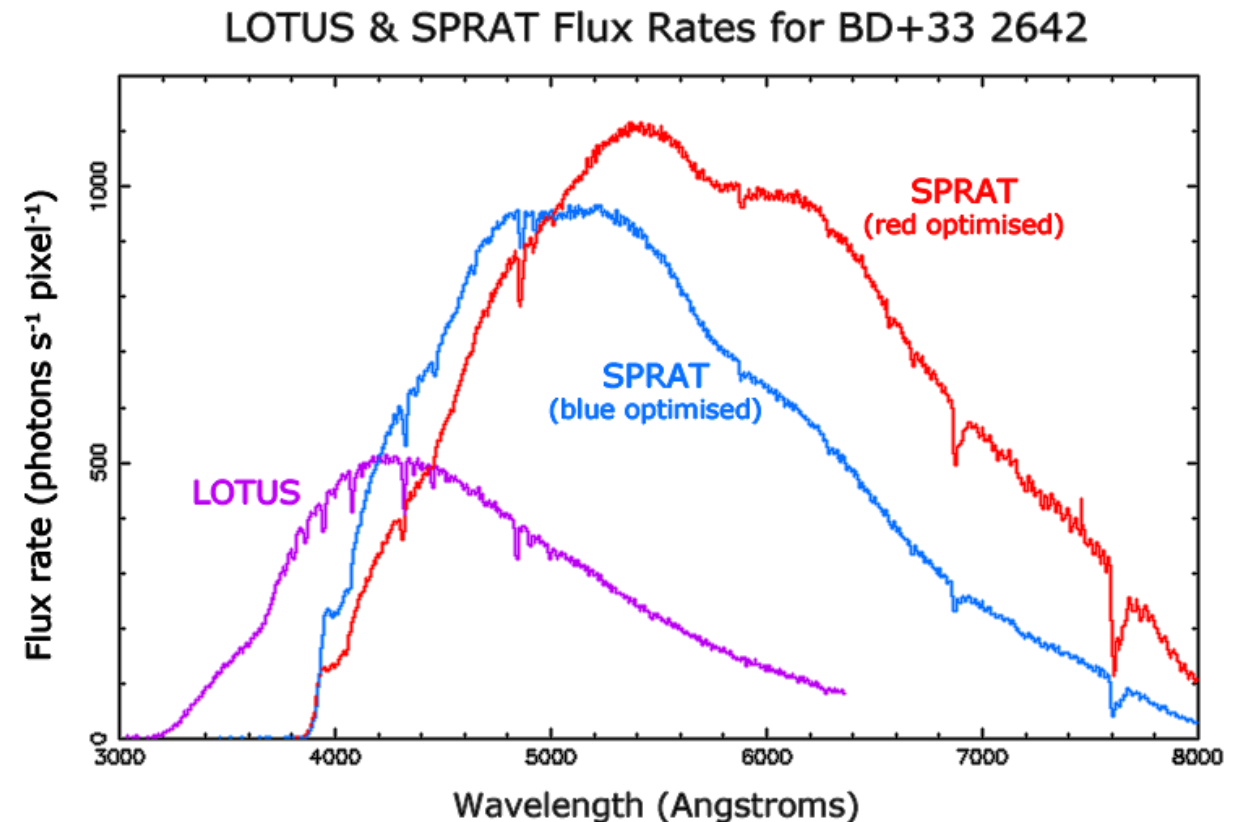
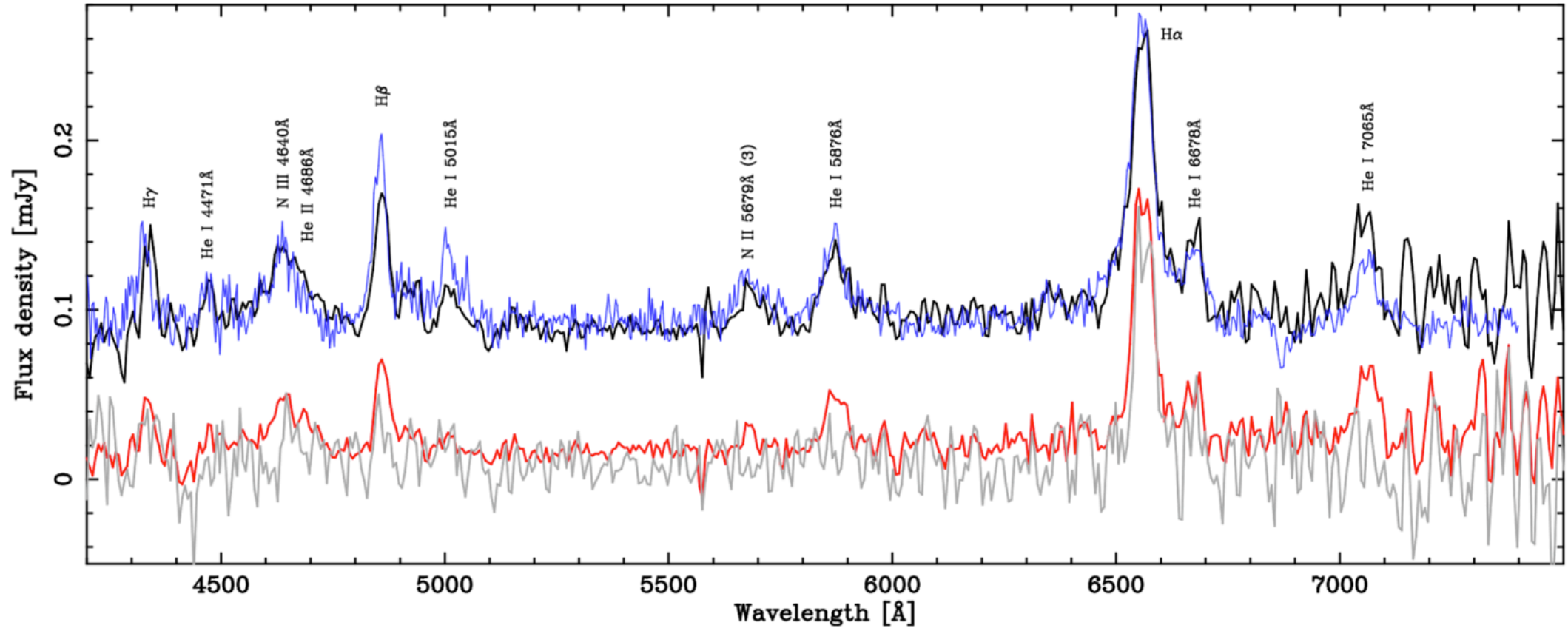


Figure 16. A plan view of the instrument identifying the optical elements. 1. Arc/Flat Calibration Angled Mirror (grey), 2. Slit (red), 3. Field Lens (blue), 4. Long-pass Filter (cyan), 5. Collimator Lens (blue), 6. 1<sup>st</sup> Prism (green), 7. VPH Grating (yellow), 8. 2<sup>nd</sup> Prism (green), 9. Camera Lens (cyan), 10. CCD (dark red)



# Sensitivity down to 20<sup>th</sup> Magnitude

A&A 580, A45 (2015)



**Fig. 3.** Liverpool Telescope SPRAT and *William Herschel* Telescope ACAM flux calibrated spectra of the 2014 eruption of M31N 2008-12a. Blue line: (WHT) taken at 2014 Oct. 3.96 UT,  $t - t_{\max} = 0.26$  days (first epoch). Black line: (LT) mean time of 2014 Oct. 4.06 UT,  $t - t_{\max} = 0.32 \pm 0.14$  days (second epoch). Red line: (LT) mean time of 2014 Oct. 5.18 UT,  $t - t_{\max} = 1.44 \pm 0.18$  days (third epoch). Grey line: (LT) taken at 2014 Oct. 5.91 UT,  $t - t_{\max} = 2.17 \pm 0.01$  days (fourth epoch).

# Gaia Alerts – SPRAT Observing Programme

- Attempt to construct an un-biased sample of spectroscopic follow-up of Gaia alerts. Selection criteria:
  - Gaia Mag < 18.5
  - Declination > -30 degrees
  - Time Since Alert < 14 days
- Also observed a similar sample of ASASSN transients for comparison
- First however, an analysis of the entire Gaia alert stream...

# Rates Increasing. (Day 0 = 27<sup>th</sup> July 2014)

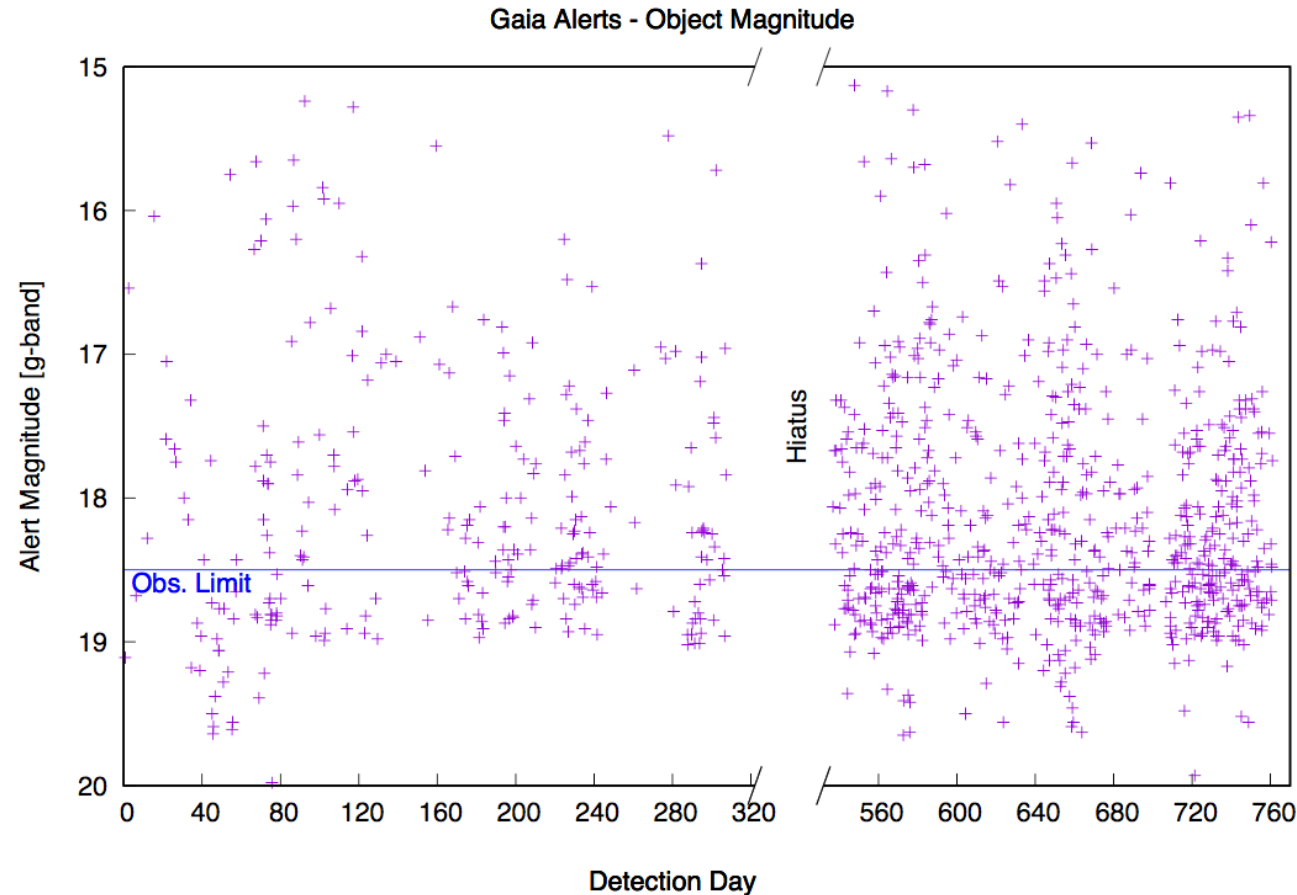


Figure 4.4: Gaia Photometric Survey alerting magnitudes. Spectroscopic observation were scheduled for alert objects brighter than 18.5 magnitude. It was noted that at day 74 of the validation phase alerts were limited  $< 19.0$  although this selection was removed post-hiatus.

# Now providing all sky coverage

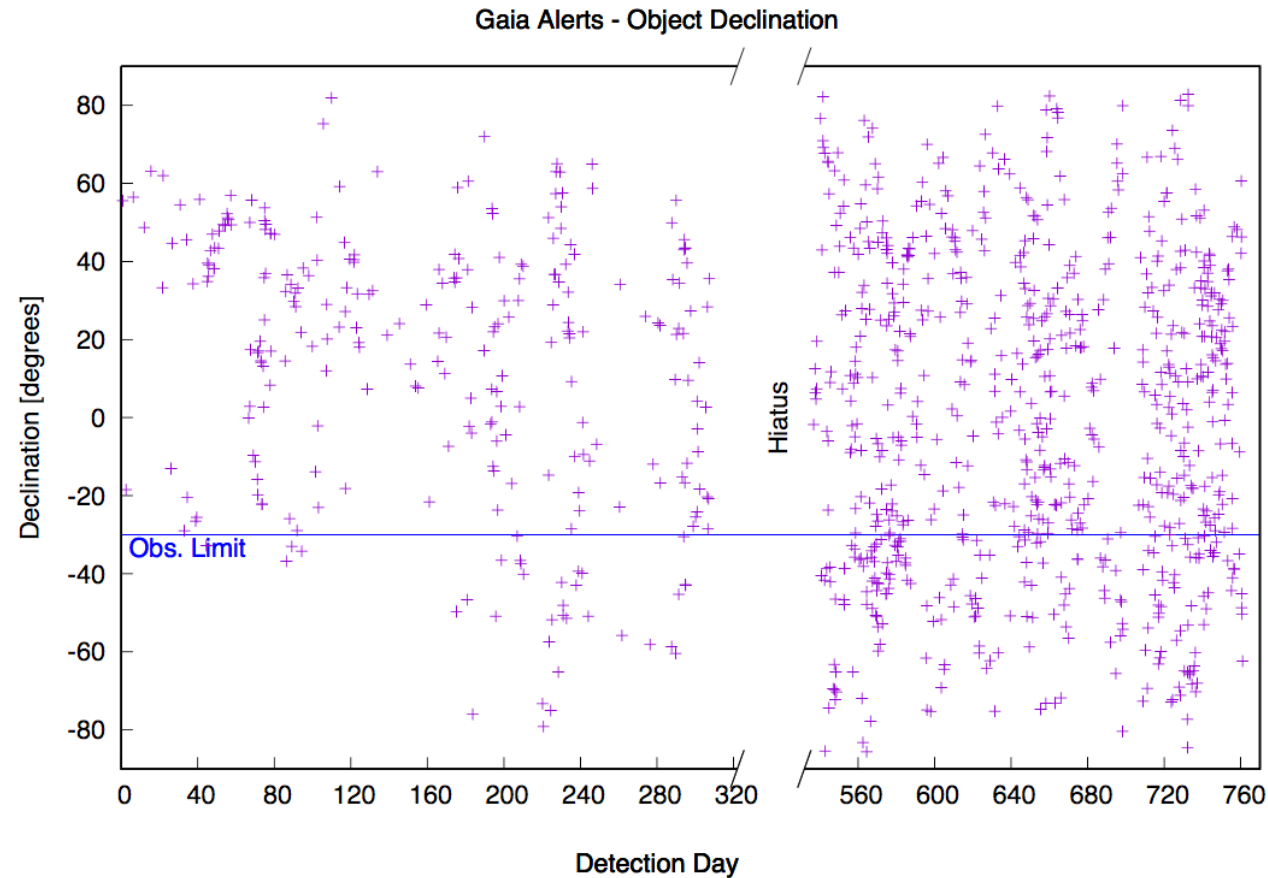


Figure 4.5: Gaia Photometric Survey object declination. The declination observing limit was set to  $-30^{\circ}$ . Some objects above the limit were not observable because of RA. The initial validation phase exhibited a northern hemisphere bias with few objects below  $-30^{\circ}$ .

# Publication delay is improving (with “blips”)

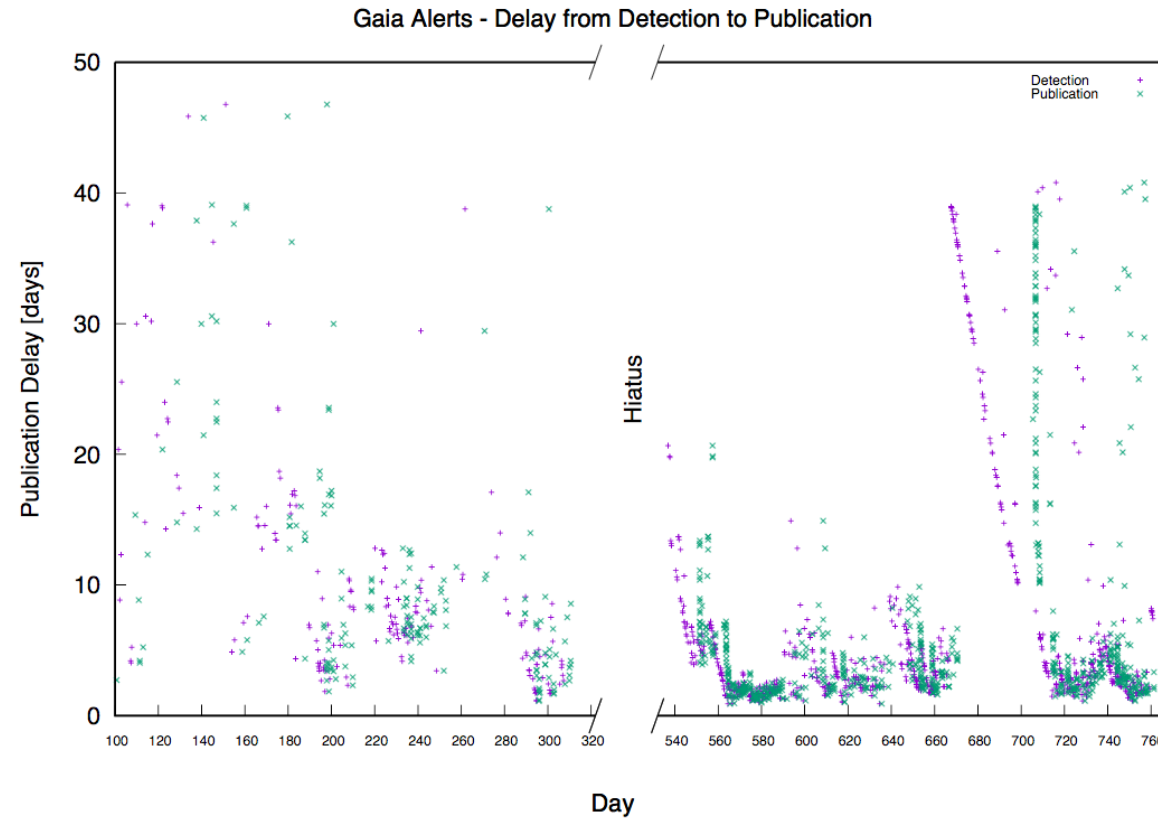


Figure 4.6: The delay between detection and publication of Gaia alerts often exhibited an episodic pattern even post hiatus. The accumulation of detections is shown by a linear decrease in delay (purple) preceding a vertical set of published alerts points (green). The alert list only began providing both the detection and publish dates after 2014-11-07 (day 104). Delays > 50 days are not shown.



# Gaia Alerts – completeness limit $\sim 19^{\text{th}}$ mag

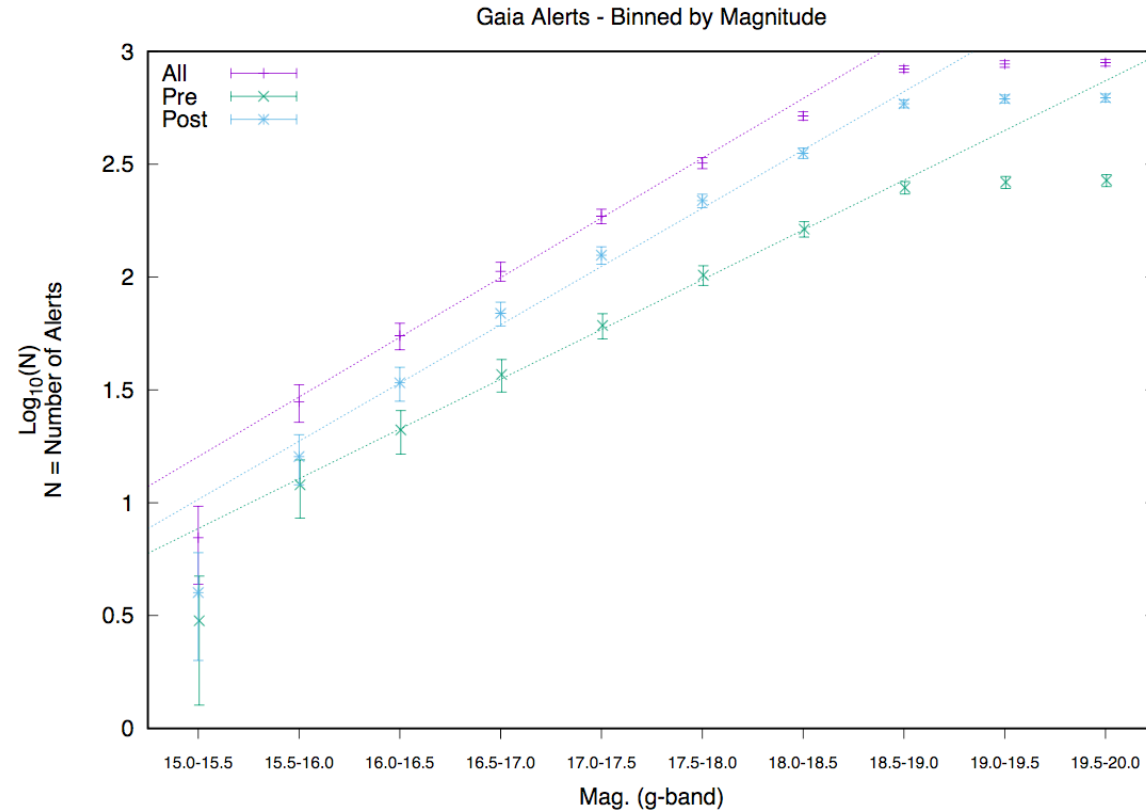


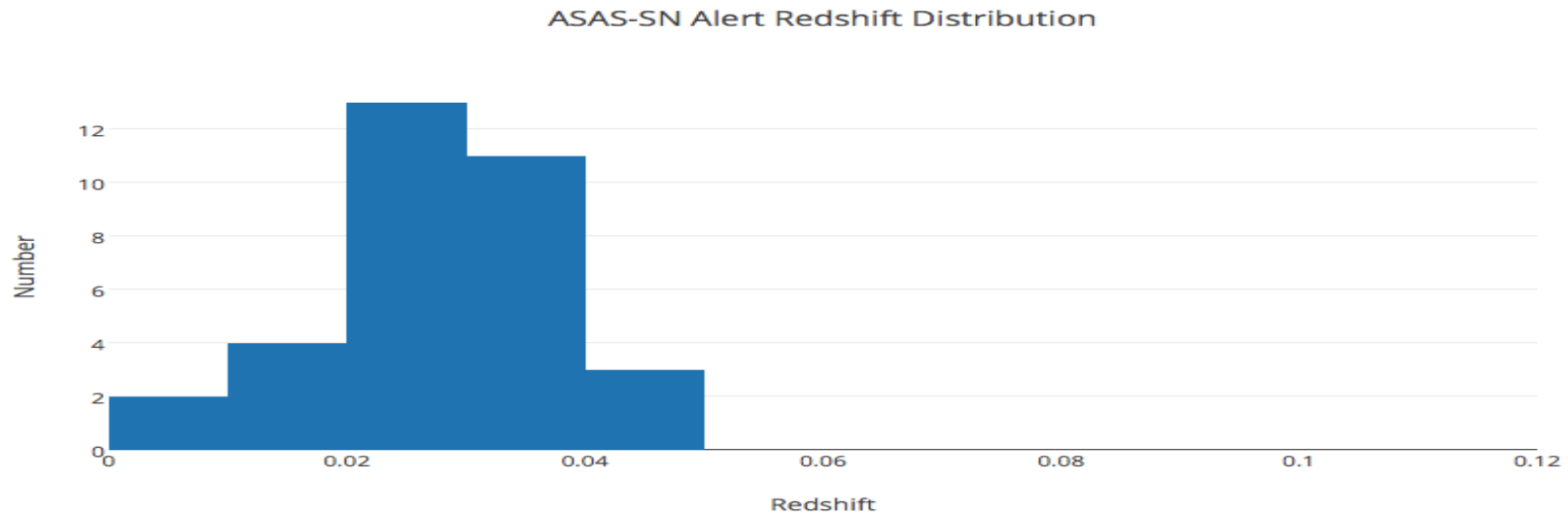
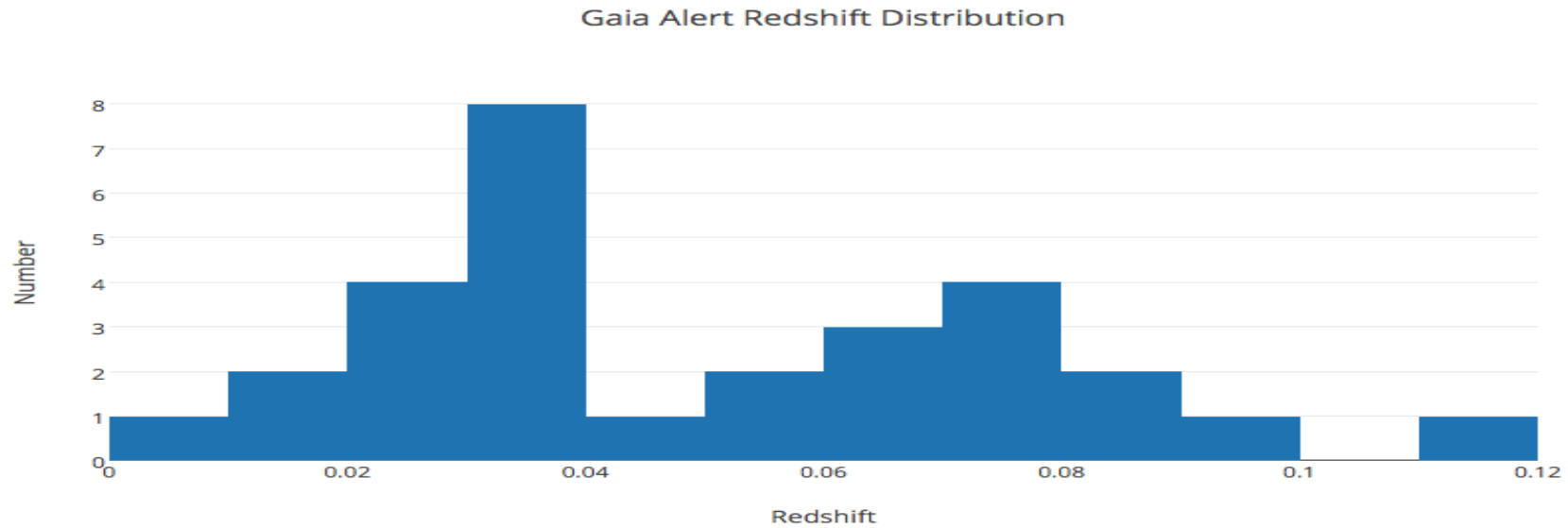
Figure 4.8: The number of alerts generated per 0.5 magnitude interval. The first 8 points show a linear relationship for all 3 sample sets and the dotted lines are fitted to these points. Fainter than magnitude 19.0 and the LogN/LogS relationship breaks down indicating the data is incomplete.

# What did we manage to observe with SPRAT?

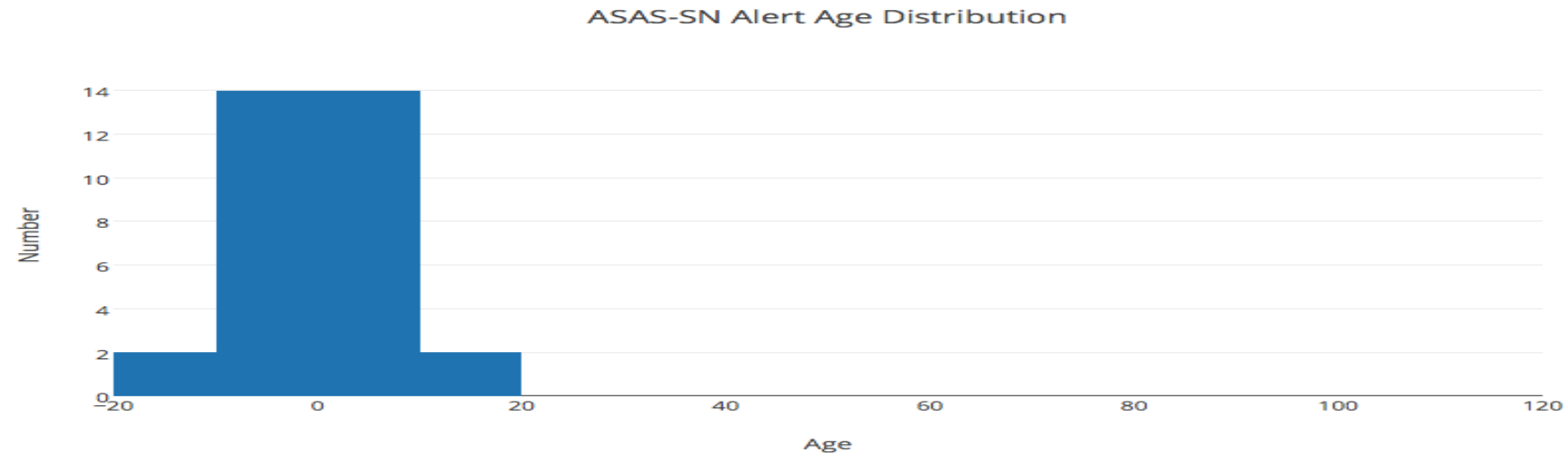
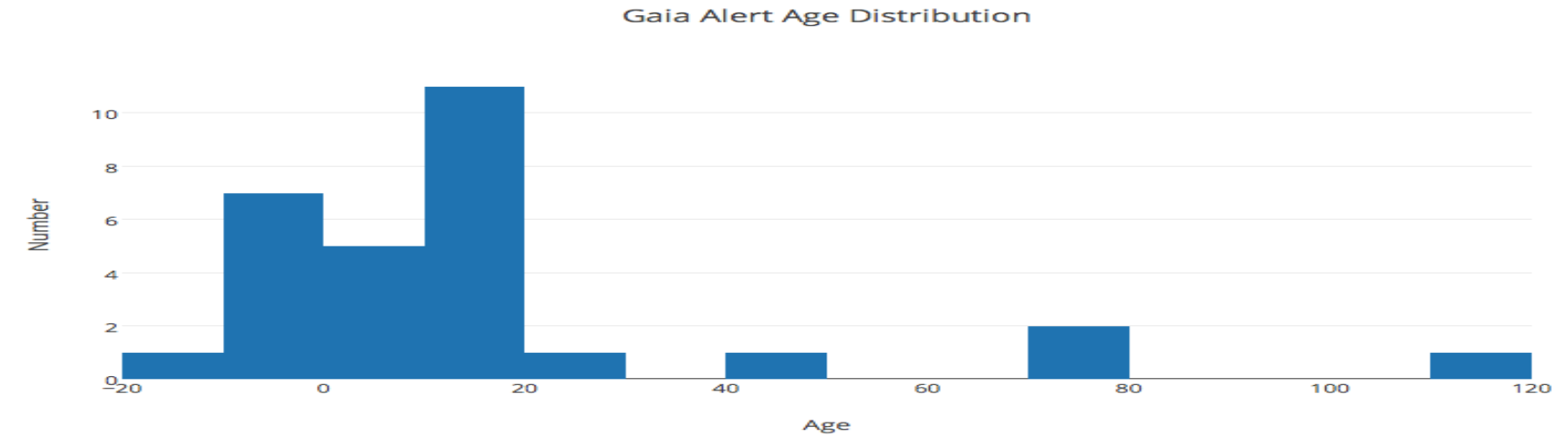
- 37 Gaia Transients.
  - 2 unclassified (noisy/poor spectra)
  - 3 normal galaxy spectra (transient faded/missed?)
  - 1 AGN
  - 31 Supernovae
- 37 ASAS-SN Transients (mainly during Hiatus).
  - 2 unclassified (noisy/poor spectra)
  - 1 flare star
  - 1 normal galaxy spectra (transient faded/missed?)
  - 1 AGN
  - 32 Supernova

Very similar overall distributions of object types.

# Gaia Alerts go to higher redshift

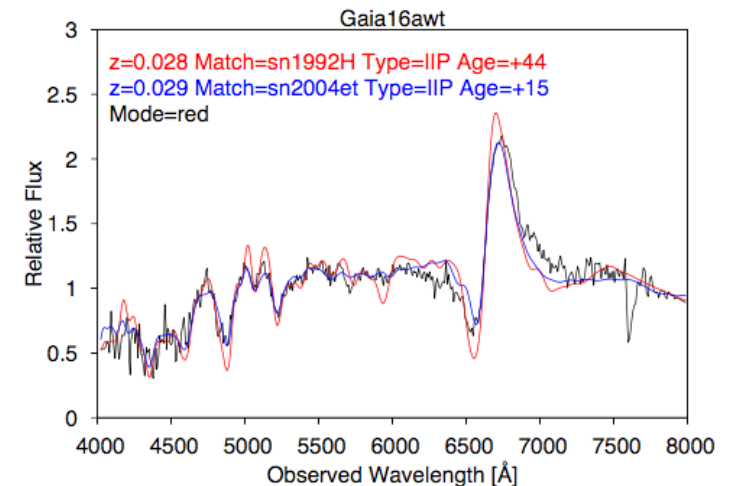
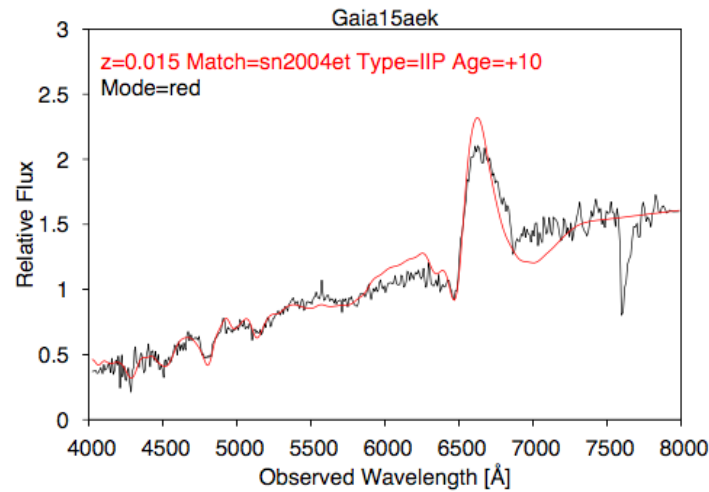
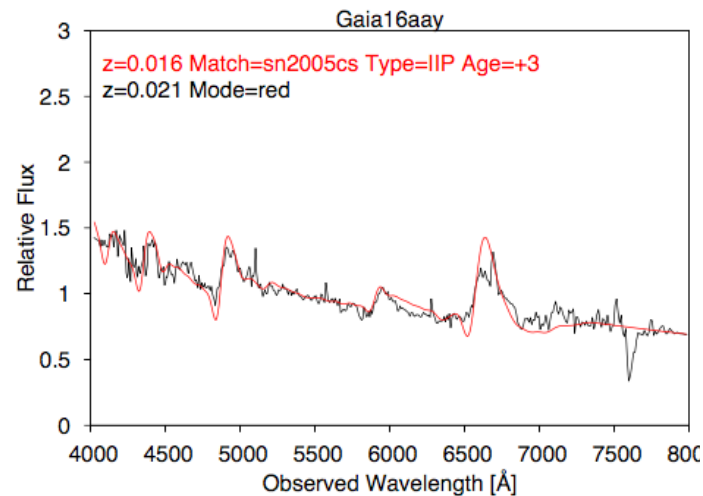


# Gaia Alerts older (although still some young)



# Gaia finds more core-collapse SNe

- Gaia (SNe Ib/Ic/II) = 32%. ASAS-SN (Sne Ib/Ic/II) = 12%



Age

# Summary

- Gaia Science Alerts are providing an interesting source of relatively bright supernovae transients out to redshift  $\sim 0.1$ .
- A reasonable fraction ( $\sim 25\%$ ) are pre maximum.
- A reasonable fraction ( $\sim 30\%$ ) are core collapse.
- A spectrograph such as SPRAT on a 2.0m class telescope is sufficient to classify and follow them up.
- Are you interested in a (simplified) low-cost (maybe 20-50k Euro?) copy of SPRAT for your 1-3 m class telescope? We have some OPTICON money to help develop such a thing. Please contact [i.a.steele@ljmu.ac.uk](mailto:i.a.steele@ljmu.ac.uk) if you are interested.