### What can we learn from Gaia about Dark Matter?

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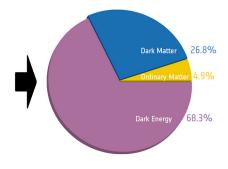


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# The missing mass

Many independent probes imply existence of Dark Matter

- Galactic rotation curves
- Cosmic Microwave Background
- Formation of Large Scale Structures
- Gravitational lensing
- Big Bang Nucleosynthesis



### What we know about DM

#### Known properties:

- Cosmological abundance
- Behaves as non-relativistic particles
- Must be stable or extremely long-lived
- It must consist of non-Standard Model particle(s)
- Strong constraints on coupling to ordinary matter

Despite great theoretical and experimental efforts not much is known about DM!



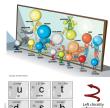
### What we don't know about DM

### Some pressing questions regarding DM

- Production mechanism?
- Single particle or multiple components?
- Distribution on small scales?

### Myriad of theoretical proposals

- Supersymmetry
- Sterile neutrinos
- Primordial Black Holes
- ...





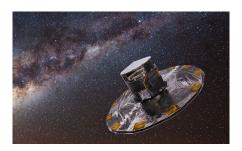


⇒ Need new observations to rule out or confirm models

# Importance of Gaia for studying DM

Gaia mission is expected to provide answers to many fundamental questions regarding DM by addressing:

- Distribution of DM within Milky Way
- Provide information regarding DM substructure
- Constrain MACHOs



# Distribution of DM in Milky Way

By mapping **proper motions of stars** one can reconstruct the total gravitational potential and hence also DM distribution in our galaxy

Knowledge DM distribution is crucial for DM searches

• Direct Detection (of DM scattering with nuclei):

$$\frac{\mathrm{d}R}{\mathrm{d}E}(E,t) = \frac{\rho_{\mathrm{DM}}(r_{\odot})}{m_{\chi}m_{A}} \int \mathrm{d}^{3}v \, \frac{\mathrm{d}\sigma}{\mathrm{d}E}(E,v) \cdot f(\vec{v},t) \cdot v$$

• Indirect Detection (of DM decay or annihilation products):

$$\frac{\mathrm{d}\Phi}{\mathrm{d}E} = \frac{\sigma_0}{8\pi m_{\nu}^2} \frac{\mathrm{d}N}{\mathrm{d}E} \int \mathrm{d}\Omega \int \mathrm{d}\ell \; \rho_{\scriptscriptstyle \mathrm{DM}}^2(\vec{r})$$

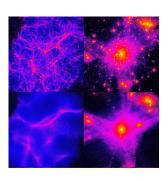
## DM substructure

By measuring **proper motions of stars** one can study the amount of DM substructure

- Disruption of Globular Clusters' tidal streams
- Stellar wakes

This could provide crucial information regarding DM nature:

- Significance of self-interactions
- Is DM "warm" or "cold"
- Might provide insights into production mechanism
- Consistency check for simulations



### MACHOs

Significant amount (if not all) DM could be in form of MAssive Compact Halo Objects

### Most popular MACHOs:

- Primordial Black Holes
- Neutron stars
- Brown dwarfs
- Unassociated planes



Primordial Black Holes are considered as most probable MACHOs candidate

### Primordial Black Holes - current status

Shrinking gap for fraction of Primordial Black Holes

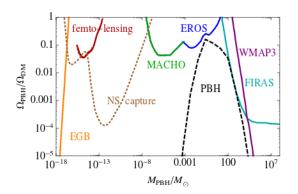


Figure: Clesse, Sébastien et al. Phys.Rev. D92 (2015) no.2, 023524

### Conclusions

Gaia has the ability to provide new observations that could be crucial for our understanding of DM

Valuable observations for many DM related research programs

- ullet DM distribution within Milky Way o Direct and Indirect Detection experiments
- ullet DM power spectrum down to the smallest scales ightarrow fundamental properties of DM
- Search for exotic components in form of MACHOs

Exciting era for DM physics!