### Gaia Science Alerts: Status



### Simon Hodgkin

E. Breedt, A. Delgado, D.L. Harrison, M. van Leeuwen, G.Rixon, T. Wevers, A. Yoldas, N. Blagorodnova, S. Koposov (IoA Cambridge), N. Ihanec, K. Kruszyńska, K.A. Rybicki, Ł. Wyrzykowski (Warsaw UniversityAstronomical Observatory), Z. Kostrzewa-Rutkowska (Leiden Observatory), D. Eappachen (SRON/RU), G. Marton (Konkoly Observatory) **@SpacemanSimon** http://gsaweb.ast.cam.ac.uk/alerts @gaia\_alerts

Galactic Longitude





Simon Hodgkin, IoA, Cambridge, UK

### Initial Data Treatment

Candidate matching: 3 sources (red), 180 observations (blue)



Fabricius et al. 2016



# Alert Detection: Daily

brighter than 19 mag

- transients
- supernovae

new

- novae
- DNe
- TDE
- AGN flares
- GRB OT
- M-dwarf flares





# detection, filtering, eyeballing



guy







diana



arancha







- numbers of transits processed by AlertPipe
- numbers of alert 2. detections
- numbers of alerts 3. presented to eyeballers after additional (mostly environmental filtering)
- numbers of alerts 4. published from each run.





### Filters

- 1. Most false positives are environmental in origin, examples:
  - 1. In wings of bright star
  - 2. planet or bright star close to pointing
  - 3. Near neighbour
- 2. But we also classify/reject:
  - Known SSOs 1.
  - 2. Known DR2 variable stars
  - 3. High amplitude red variables





### Filters

- 1. Most false positives are environmental in origin, examples:
  - 1. In wings of bright star
  - 2. planet or bright star close to pointing
  - 3. Near neighbour
- 2. But we also classify/reject:
  - Known SSOs
  - 2. Known DR2 variable stars
  - 3. High amplitude red variables







### 14,961 Alerts since Nov 2014



- Initial commissioning phase, spanning almost 300 days
- Density map avoiding crowded regions used for few months
- Environment Filters and Eyeballing App.
- Rock and GDR2 variables
- Skew-VonNewumann detector introduced



## **Precise Photometry**

- Photometric One-Day 0.10 Calibration 0.08
- Precision per-transit
- $\sigma \sim 1\%$  at G=15
- $\sigma \sim 3\%$  at G=19
- Updated ~monthly
- DPCI (Paul Osborne)



- STD (PODC) 0.04
  - 0.02
  - 0.00

### **Transits of Alerts**



### Accurate Astrometry GSA vs ZTF/PS1/ATLAS/OGLE GSA vs GDR2



- TNS AstroNote 2019-15
- ...we used the transient reports by Gaia as "ground truth"



'GaiaAlerts' vs 'Pan-STARRS1' (bin width = 0.05, points = 667, median = 0.118, mean = 0.149, xaxis range = [0.0, 2.0] )



'GaiaAlerts' vs 'ATLAS' (bin width = 0.05, points = 1063, median = 0.257, mean = 0.335, xaxis range = [0.0, 2.0] )



'GaiaAlerts' vs 'OGLE' (bin width = 0.05, points = 22, median = 0.209, mean = 0.236, xaxis range = [0.0, 2.0] )



0.3



Simon Hodgkin, IoA, Cambridge, UK











### Classification

- almost 25 per cent of alerts classified.
- These classifications were obtained from numerous sources (including Simbad, ADS, TNS, ATels)
- SN dominate: due to dedicated spectroscopic follow-up programmes (PESSTO, NUTS, SED Machine, ZTF BTS)
- Thus the classification fractions should not be extended to the entire sample of GSA events.



### Classification

### Classified



Majority of unclassified alerts are Galactic in origin, and thus the statistics presented in previous figure are not reflective of the true breakdown of the GSA transient classes.

Simon Hodgkin, IoA, Cambridge, UK

### Unclassified

) 2

### Classification



**Left**: CMD of GDR2 counterparts to Gaia Alerts (stars), overlaid on a clean 100pc sample (grey background). Classified CVs are blue and YSOs are red. **Right**: probability map of a binary SVM classifier for 1815 unknown alerts in GDR2. Blue regions indicate CV parameter space, red regions YSOs. The white line is the decision boundary; training samples are shown as coloured stars, while new classifications (with P>0.95) are shown as magenta/pink circles. 727 (838) sources as CVs for P>0.95 (P>0.90), while 198 (410) new YSOs are classified.

Simon Hodgkin, IoA, Cambridge, UK



•	Minimum purity of the published alerts as a function of Galactic latitude	0.8	
		0.7	
•	An alert considered real if there		
	exists some ancillary information	0.6	
	concerning it, such as a	t∠	
	classification, or an ATEL, or if it was	UN 0.5	
	reported by another survey either	Ę	
	before or after the Gaia detection.	٥.4 الم	
•	We conclude that these alerts must	Σ	
	be astrophysically real and not due	0.3	
	to some known or unknown		
	systematic of the Gaia data.	0.2	
•	For $G < 17$ , the fraction of our alerts	0.1	
	confirmed by ancillary data is 0.93	0.1	_

# Purity





# neighbouring sources within 2', but outside of 0.2'', of each published alert in the AL and AC directions at the time of the alert, for  $b \ge 40^{\circ}$  (Left), and for  $|b| < 8^{\circ}$  (Right). Much larger proportion of SN detected at high latitudes (hence excess near alert position), and a lack of perturbing/contaminating neighbour sources for alerts at low Galactic latitude.

No evidence that alerts in the Galactic plane are less reliable than those at higher latitudes, nor that fainter alerts are any less reliable than brighter ones. Thus the overall purity of our published alerts is no worse than that of the subset where  $b \ge 40^{\circ}$  and Gmag < 17 and the fraction of our alerts confirmed by ancillary data is 0.93.

Simon Hodgkin, IoA, Cambridge, UK

Purity







Z

# Completeness for SNe

- We have tested the completeness of GSA
- Using SNe: classified, numerous, long-lived
- TNS is the IAU-approved host of discoveries and classifications for SNe

# Completeness: scanning law

- TNS recorded 2826
  SNe, mag<=19</li>
  (7/16-1/20)
- GSA requires detection in both FoVs within 40 dy
- This dominates GSA recovery of SNe.
- External completeness,  $C_E=0.46$



# Completeness: magnitude

- Now use sample with >=2 scans (1564 SNe)
- 1073 detected, 491 not: thus overall internal completeness  $C_I = 0.69$
- Fairly uniform with mag
- Roll-off to fainter magnitudes:
  - different survey photometry?
  - fading at time of discovery?



## Completeness: separation

- SN with >=2 scans
- xmatch vs. 2MASS
  Extended Sources
- likely mismatches for sep>~20 arcsec
- Significant drop at sep<~3 arcsec</li>
- Indicative of a reduction in the sensitivity of the NewSource detector to complex environments



# Coming Soon

- Single Transit Alerts (Kostrzewa-Rutkowska+20, Tue talk)
- Full Automation (Marton, Wed talk)
- secure their consent.

Simon Hodgkin, IoA, Cambridge, UK

GSA paper: it has been shown to DPACE, and we are addressing (mostly minor) comments. We are about to circulate to all authors to

DR3 includes alerts in Cycle 03 (25-07-2014 to 28-05-2017)

# Summary

- A series of (mostly environmental) filters reduces the detection rate from a few thousand provisional candidates to a shortlist of several tens of events per day which are eyeballed.
- Astrometry is accurate to 55 mas, photometry to better than 0.031 mag (at G=19).
- We currently publish alerts at a rate of approximately 12 events/day, ~25 per cent of them receive classifications.
- Published classifications are dominated by SNe, but this is biased by the extensive follow-up campaigns. The bulk of our unclassified alerts reside in the Galactic plane, and are therefore likely to be Galactic in origin.
- GSA suffers very low levels of contamination from false positives. We showed that the minimum purity of the survey in uncrowded regions for sources with G<17 is 93 per cent. We find no evidence of additional sources of contamination at faint magnitudes, or in crowded regions.
- Overall completeness to SNe is 46 per cent. The Gaia scanning law, combined with requirement for two detections, dominates ability to detect SNe. Completeness for SNe with>=2 scans is 79 per cent, unless within 3 arcsec of the nucleus.
- candidate YSOs.

We handle upwards of 60 million measurements per-day, searching for new sources, and sources which change significantly in brightness. Detectors work with lightcurves, and thus are also sensitive to slowly varying sources.

For those alerting sources which were included in Gaia DR2, we have shown that a simple classifier is able to distinguish between YSOs, and CVs with high reliability, resulting in 838 new candidate CVs, and 410 new

### Thanks and Acknowledgements

DPAC
Francesca De Angeli
Giuseppe Altavilla
Vasily Belokurov
Nadejda Blagorodnova
Josh Bloom
Elme Breedt
Anthony Brown
Jos de Bruijne
Patrick Burgess
Ross Burgon
Giorgia Busso
Heather Campbell
Gisella Clementini
Chris Copperwheat
Sue Cowell
Michel Dennefeld
Andrew Drake
Dafydd Wyn Evans

- Laurent Eyer
- Morgan Fraser
- Gerry Gilmore
- GREAT
- Liam Hardy
- Jorge Fernandez Hernandez
- Greg Holland
- Anna Hourihane
- Peter Jonker
- Uli Kolb
- Zuzanna Kostrzewa-Rutkowska
- Sergey Koposov
- Stuart Littlefair
- Ashish Mahabal
- Francois Mignard
- Paul Osborne
- Lovro Palaversa
- Andrzej Pigulski
- Timo Prusti

Simon Hodgkin, IoA, Cambridge, UK

- Marco Riello
- Guy Rixon
- Norbert Schartel
- lain Steele
- Rachel Street
- Frank Suess
- Lina Tomasella (and team),
- Manuel Torres
- Yiannis Tsapras
- Massimo Turatto
- Sjoert van Velzen
- Nic Walton
- Thomas Wevers
- Patricia Whitelock
- Roy Williams
- Lukasz Wyrzykowski
- all co-l's on our numerous proposals.