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Parameter	Units	Values
Parameter	Units	values

Stellar Param	eters:		
$M_* \ldots \ldots$	Mass $(M_{\odot})$	$0.946_{-0.073}^{+0.12}$	
$R_* \ldots \ldots$	Radius $(R_{\odot})$	$4.00^{+0.21}_{-0.15}$	
$R_{*,SED}$	Radius <sup>1</sup> ( $R_{\odot}$ )	$4.12_{-0.24}^{+0.28}$	
$L_* \dots$	Luminosity $(L_{\odot})$	$7.49^{+1.2}_{-0.99}$	
$F_{Bol}$	Bolometric Flux (cgs)	0.000000000291+0.00000000000000000000000000000000000	
$\rho_* \dots$	Density (cgs)	$0.02059\substack{+0.0015\\-0.00048}$	
log <i>g</i>	Surface gravity (cgs)	$3.210^{+0.031}_{-0.018}$	
$T_{\rm eff}$	Effective Temperature (K)	$4770_{-150}^{+140}$	
$T_{\rm eff,SED}$	Effective Temperature <sup>1</sup> (K)	$4720_{-160}^{+130}$	
[Fe/H]	Metallicity (dex)	$-0.20^{+0.23}_{-0.38}$	
$[Fe/H]_0$ .	Initial Metallicity <sup>2</sup>	$-0.24^{+0.22}_{-0.36}$	
<i>Age</i>	Age (Gyr)	$11.3^{+1.9}_{-3.5}$	
<i>EEP</i>	Equal Evolutionary Phase <sup>3</sup>	$491.8^{+1.9}_{-1.7}$	
$A_V \dots$	V-band extinction (mag)	$1.83^{+0.14}_{-0.21}$	
$\sigma_{SED} \dots$	SED photometry error scaling	$11.7^{+1.8}_{-1.5}$	
ω	Parallax (mas)	$0.347^{+0.023}_{-0.025}$	
<i>d</i>	Distance (pc)	$2880^{+220}_{-180}$	
Planetary Par	ameters:	b	

<i>P</i>	Period (days)	$1.0213727^{+0.0000084}_{-0.0000097}$
$R_P \ldots \ldots$	Radius ( <i>R</i> <sub>J</sub> )	$1.407\substack{+0.10\\-0.069}$
$M_P \ldots$	$Mass^4 (M_J)$	$1.16^{+14}_{-0.75}$
$T_C \ldots \ldots$	Time of conjunction <sup>5</sup> (BJD <sub>TDB</sub> )	$2455376.673^{+0.014}_{-0.013}$
$T_T \ldots \ldots$	Time of minimum projected separation <sup>6</sup> (BJD <sub>TDB</sub> )	$2455376.673^{+0.014}_{-0.013}$
$T_0$	Optimal conjunction Time <sup>7</sup> (BJD <sub>TDB</sub> )	$2456398.0454^{+0.0094}_{-0.0095}$
<i>a</i>	Semi-major axis (AU)	$0.01956^{+0.00098}_{-0.00055}$
<i>i</i>	Inclination (Degrees)	$78.8_{-12}^{+7.8}$
$T_{eq}$	Equilibrium temperature <sup>8</sup> (K)	$3280^{+100}_{-110}$
$ au_{ m circ} \dots$	Tidal circularization timescale (Gyr)	$0.00072^{+0.013}_{-0.00050}$
<i>K</i>	RV semi-amplitude <sup>4</sup> (m/s)	$240^{+3000}_{-160}$
$R_P/R_*$	Radius of planet in stellar radii	$0.0362\substack{+0.0013\\-0.0012}$
$a/R_*\ldots$	Semi-major axis in stellar radii	$1.0456^{+0.025}_{-0.0071}$
$\delta$	$\left(R_P/R_*\right)^2$	$0.001312\substack{+0.000095\\-0.000083}$
$\delta_{\mathrm{I}}$	Transit depth in I (fraction)	$0.00165^{+0.00011}_{-0.00010}$
$\delta_{\mathrm{V}} \dots \dots$	Transit depth in V (fraction)	$0.00192\substack{+0.00016\\-0.00015}$
$ au \dots$	Ingress/egress transit duration (days)	$0.045^{+0.010}_{-0.012}$
$T_{14}$	Total transit duration (days)	$0.465^{+0.024}_{-0.040}$

Table 1 continued on next page

Table 1 (continued)

Parameter	Units		Value	S
$T_{FWHM}$	FWHM transit duration (days)	$0.420^{+0.016}_{-0.031}$		
<i>b</i>	Transit Impact parameter	$0.21\substack{+0.21\\-0.14}$		
$\delta_{S,2.5\mu m}$	Blackbody eclipse depth at $2.5\mu m$ (ppm)	$642_{-44}^{+50}$		
$\delta_{S,5.0\mu m}$	Blackbody eclipse depth at $5.0\mu m$ (ppm)	$773^{+57}_{-50}$		
$\delta_{S,7.5\mu m}$	Blackbody eclipse depth at $7.5\mu m$ (ppm)	$816_{-52}^{+60}$		
$\rho_P \ldots \ldots$	Density <sup>4</sup> (cgs)	$0.54_{-0.36}^{+8.2}$		
$logg_P \dots$	Surface gravity <sup>4</sup>	$3.17^{+1.2}_{-0.48}$		
$\Theta \dots$	Safronov Number	$0.035^{+0.48}_{-0.023}$		
$\langle F \rangle \dots$	Incident Flux $(10^9 \text{ erg s}^{-1} \text{ cm}^{-2}) \dots$	$26.3^{+3.5}_{-3.3}$		
$T_P \ldots \ldots$	Time of Periastron (BJD <sub>TDB</sub> )	$2455376.673_{-0.013}^{+0.014}$		
$T_S \ldots \ldots$	Time of eclipse (BJD <sub>TDB</sub> )	$2455377.183^{+0.014}_{-0.013}$		
$T_A \ldots \ldots$	Time of Ascending Node (BJD <sub>TDB</sub> )	$2455377.439_{-0.013}^{+0.014}$		
$T_D \ldots \ldots$	Time of Descending Node (BJD <sub>TDB</sub> )	$2455376.928^{+0.014}_{-0.013}$		
$V_c/V_e\ldots$		1.00		
$M_P \sin i$ .	Minimum mass <sup>4</sup> $(M_J)$	$1.12^{+14}_{-0.73}$		
$M_P/M_*$ .	Mass ratio <sup>4</sup>	$0.00118^{+0.015}_{-0.00077}$		
$d/R_*\ldots$	Separation at mid transit	$1.0456^{+0.025}_{-0.0071}$		
$P_T \ldots \ldots$	A priori non-grazing transit prob	$0.9214_{-0.021}^{+0.0067}$		
$P_{T,G}$	A priori transit prob	$0.9912\substack{+0.0068\\-0.023}$		
Wavelength 1	Parameters:	Ι	V	
$u_1 \ldots \ldots$	linear limb-darkening coeff	$0.440^{+0.046}_{-0.052}$	$0.683^{+0.068}_{-0.080}$	
$u_2 \ldots \ldots$	quadratic limb-darkening coeff	$0.212^{+0.043}_{-0.040}$	$0.105^{+0.066}_{-0.058}$	
Transit Parar	neters:	OGLE UT 2010-03-06 (I)	OGLE UT 2010-03-06 (V)	OGLE UT 2010-06-29
$\sigma^2 \dots$	Added Variance	$0.00003065\substack{+0.00000047\\-0.00000046}$	$0.0000319\substack{+0.000060\\-0.0000052}$	$0.00003064^{+0.00000047}_{-0.00000046}$
$F_0 \ldots \ldots$	Baseline flux	$1.000628^{+0.000063}_{-0.000064}$	$1.00049 \pm 0.00052$	$1.000629\substack{+0.000063\\-0.000064}$
See Table 3 i	n Eastman, J. et al., 2019, arXiv:1907.09480 for a deta	iled description of all paramete	rs	
<sup>1</sup> This value ig	nores the systematic error and is for reference only			
<sup>2</sup> The metallic	ity of the star at birth			
<sup>3</sup> Corresponds	to static points in a star's evolutionary history. See §2	in Dotter, A., 2016, ApJS, 222,	8	
<sup>4</sup> Uses measur	ed radius and estimated mass from Chen. J., & Kinning	g, D. 2017, ApJ. 834. 17		
5 <sub>Time of coni</sub>	unction is commonly reported as the "transit time"			
rime or conj	unedon is commonly reported as the transit time			

<sup>6</sup>Time of minimum projected separation is a more correct "transit time"

<sup>7</sup>Optimal time of conjunction minimizes the covariance between  $T_C$  and Period

<sup>8</sup>Assumes no albedo and perfect redistribution