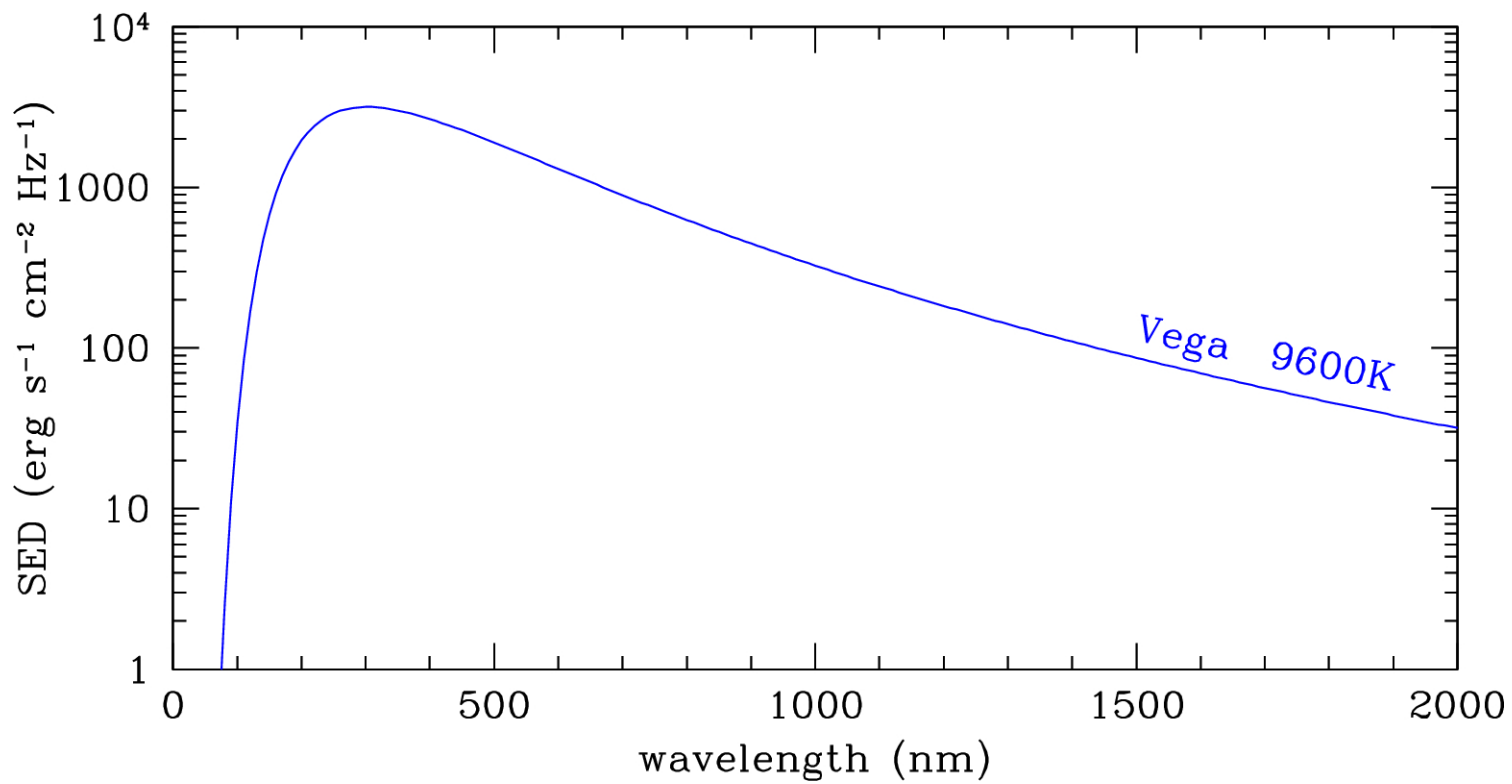


Analiza Widmowa Obiektów

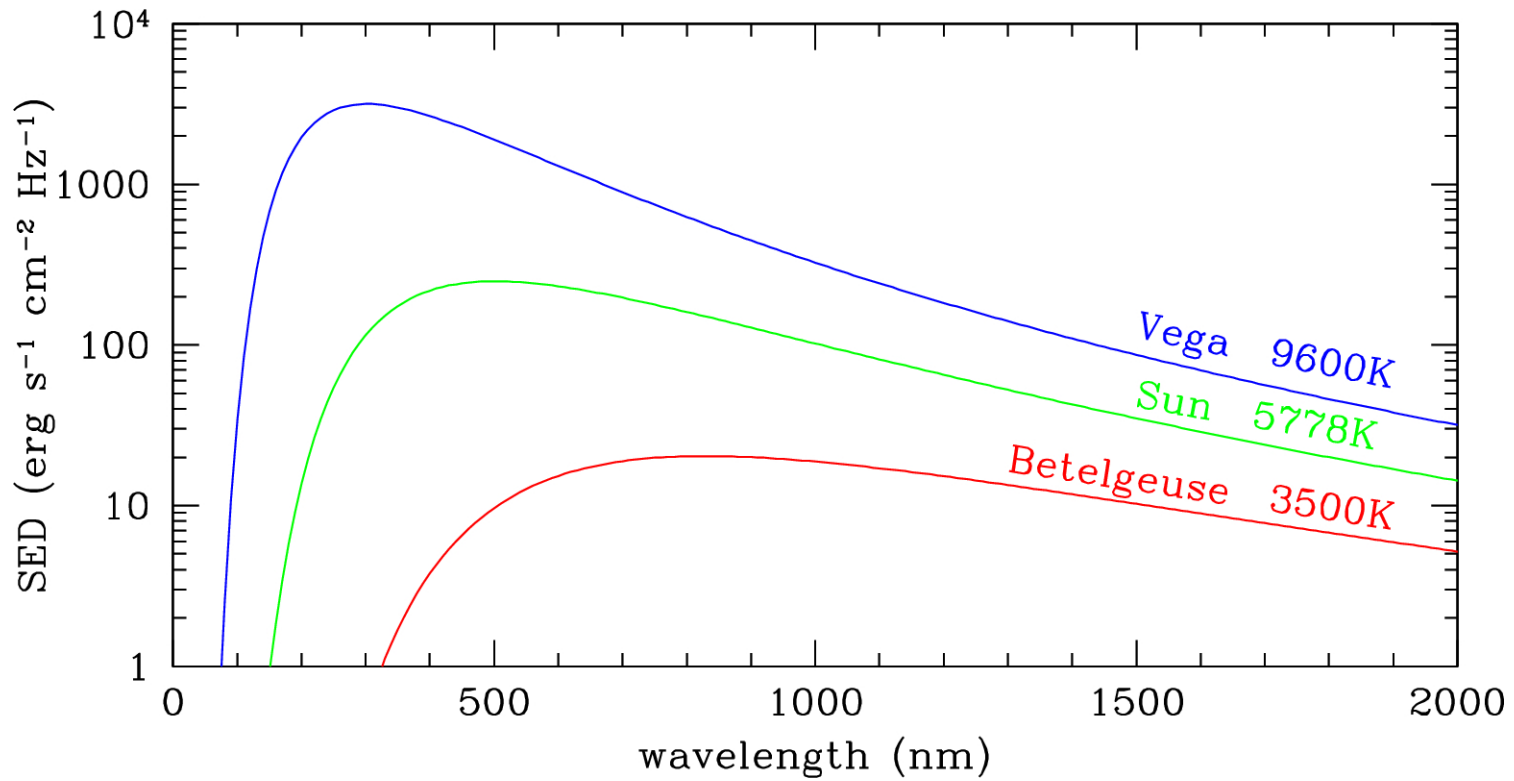
Szymon Kozłowski

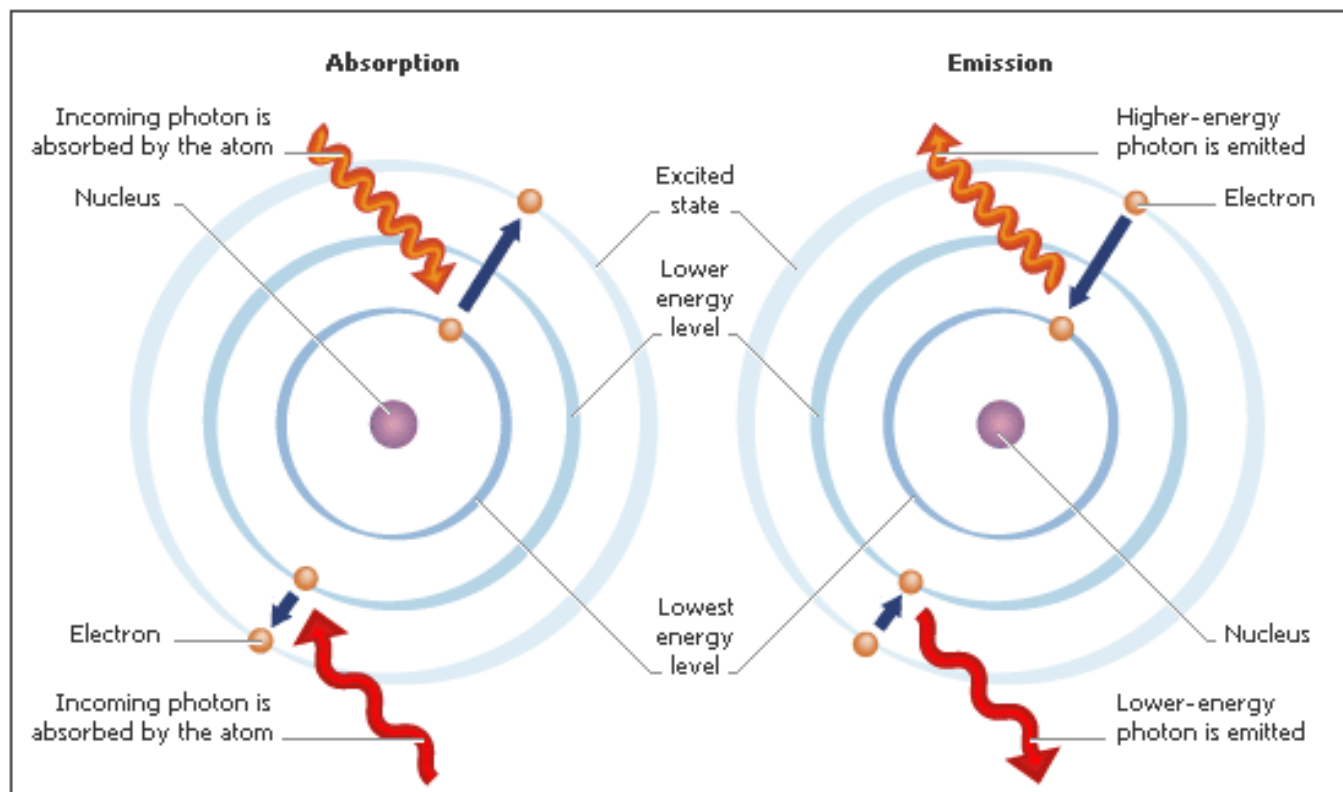


Widmo ciągłe - kontinuum

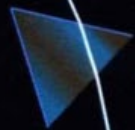


Widmo ciągłe - kontinuum





Hot blackbody



Prism

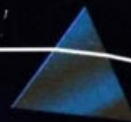


a

Continuous spectrum

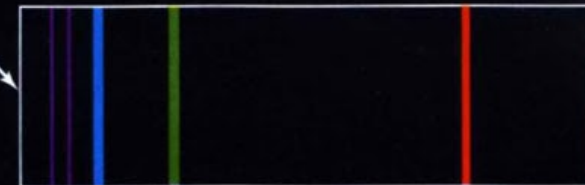
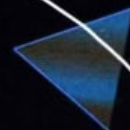
Cloud of cooler gas

Prism



b

Absorption line spectrum



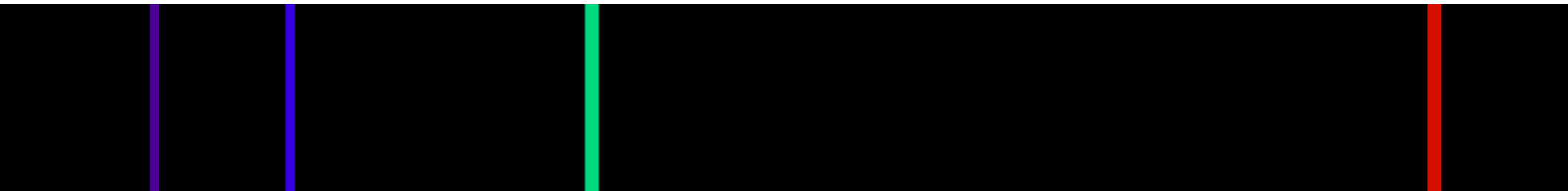
c

Emission line spectrum

Widmo obiektów



widmo ciągłe – kontinuum (continuum)

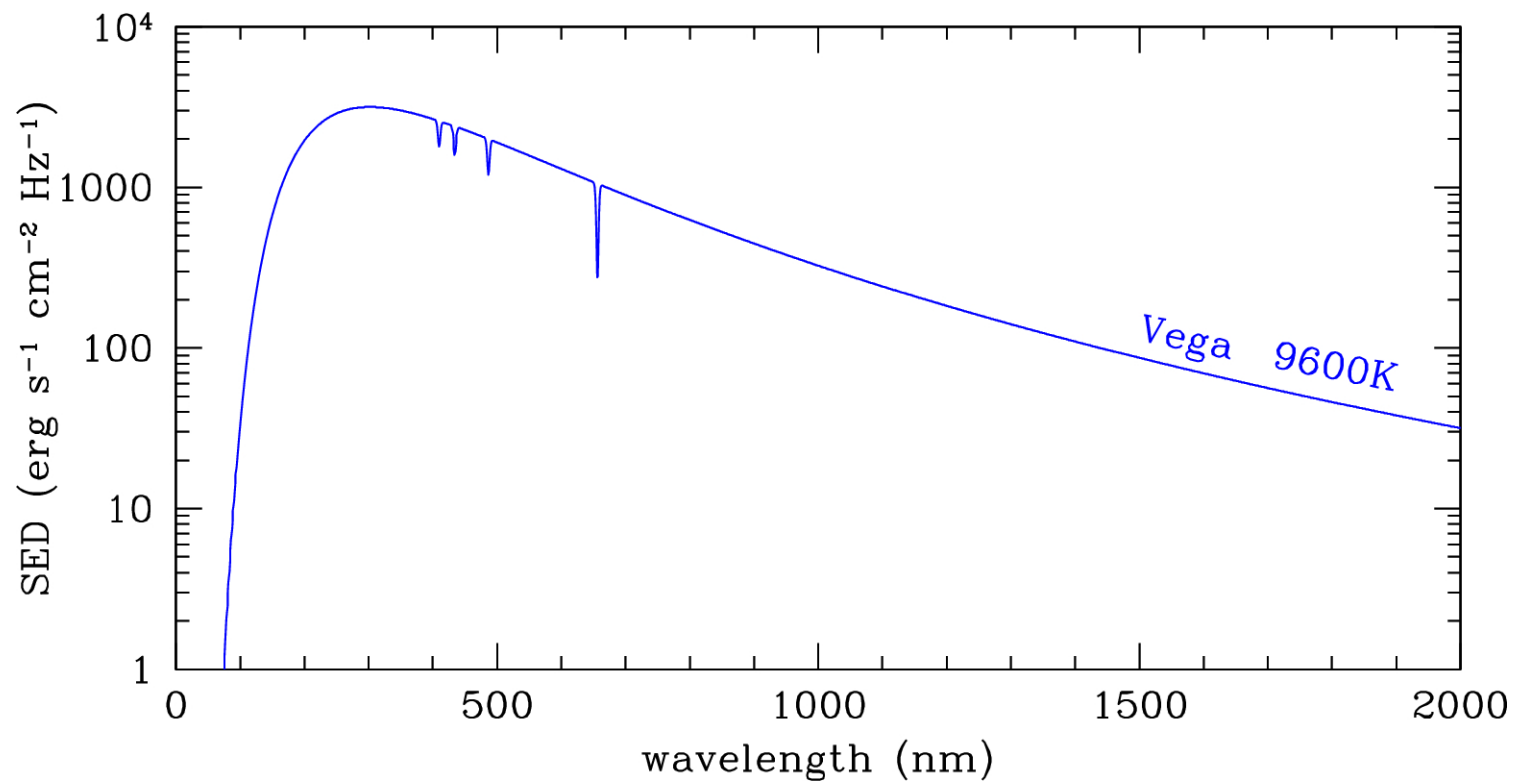


linie emisyjne (emission lines)

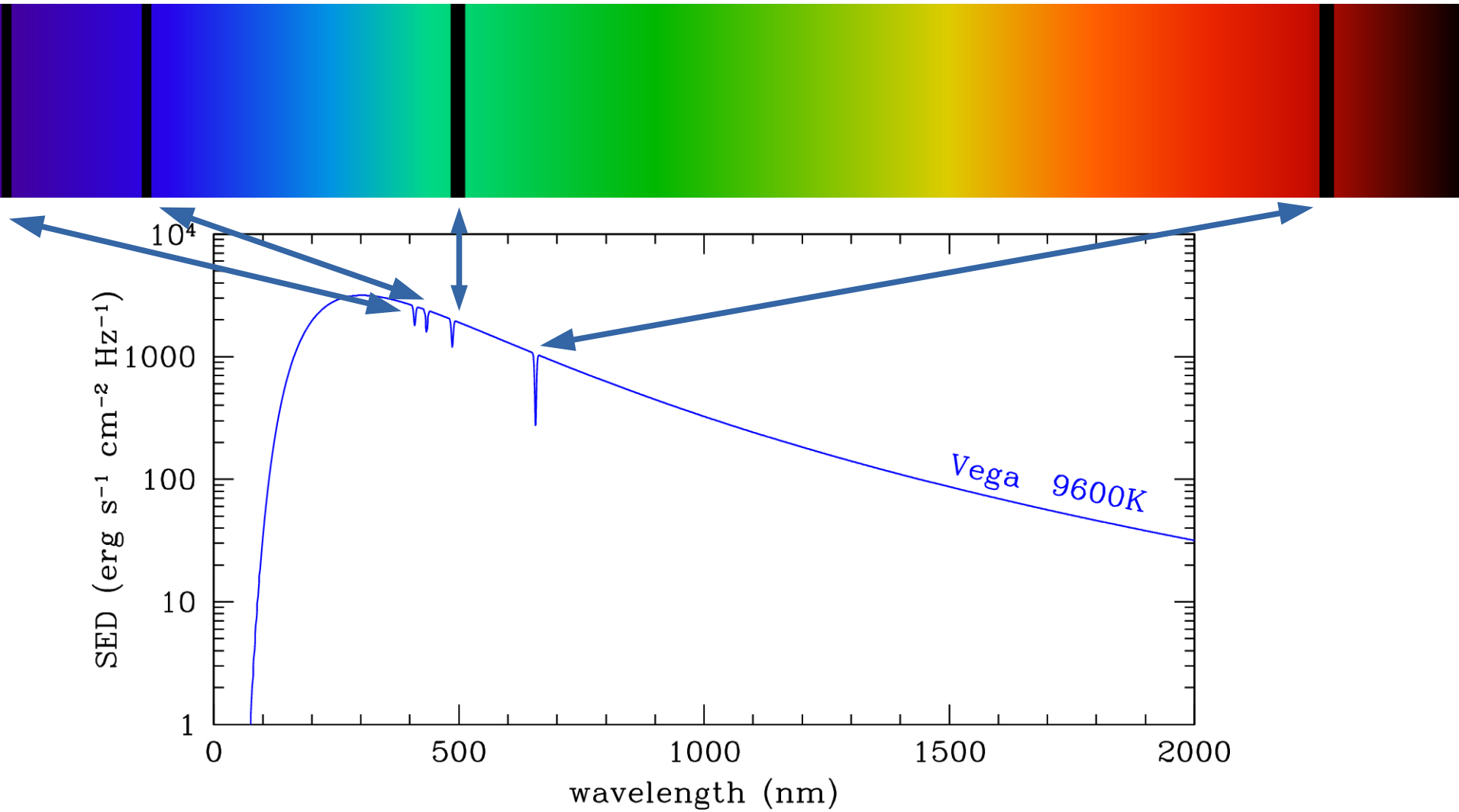


linie absorpcyjne (absorption lines)

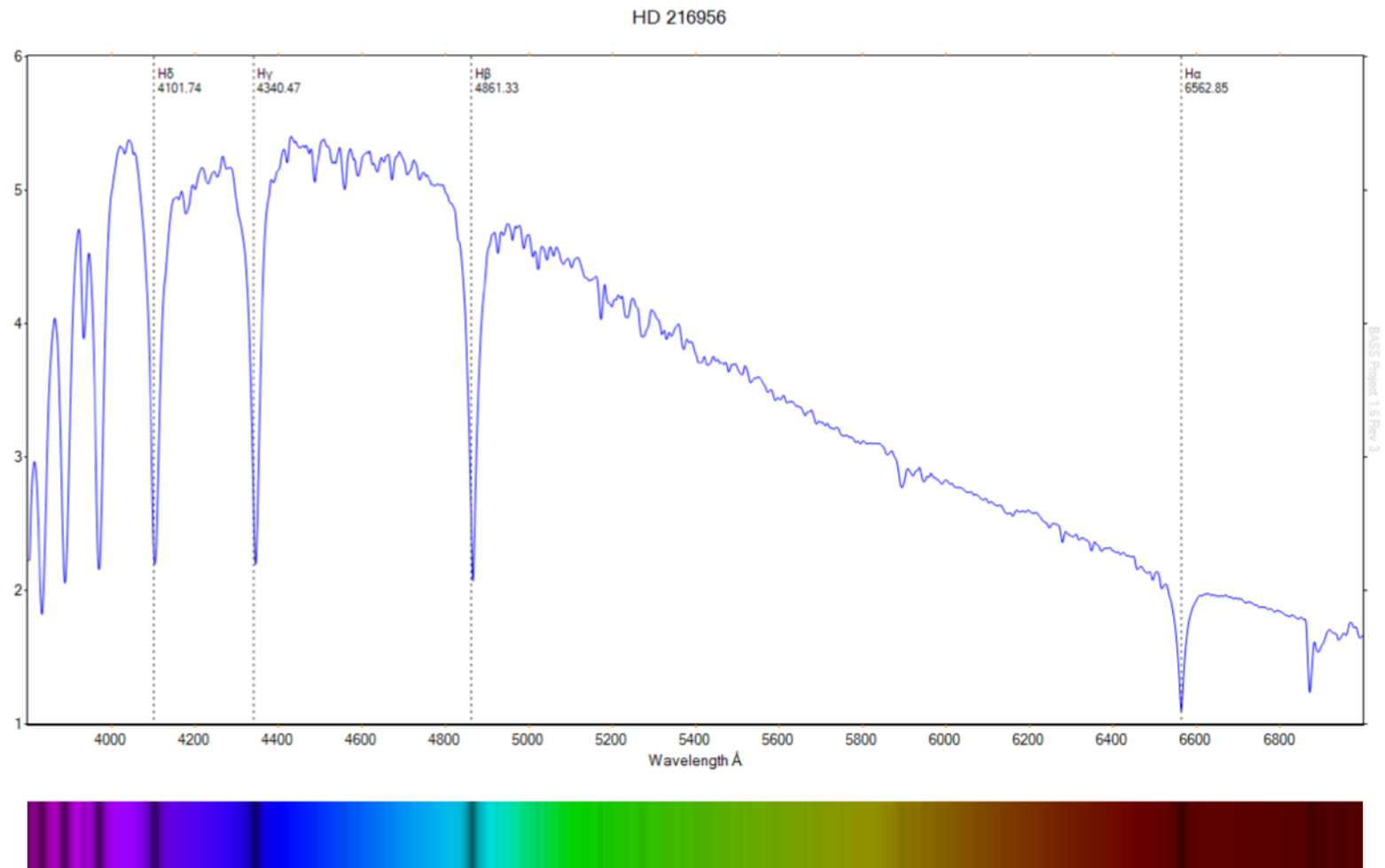
Linie absorpcyjne



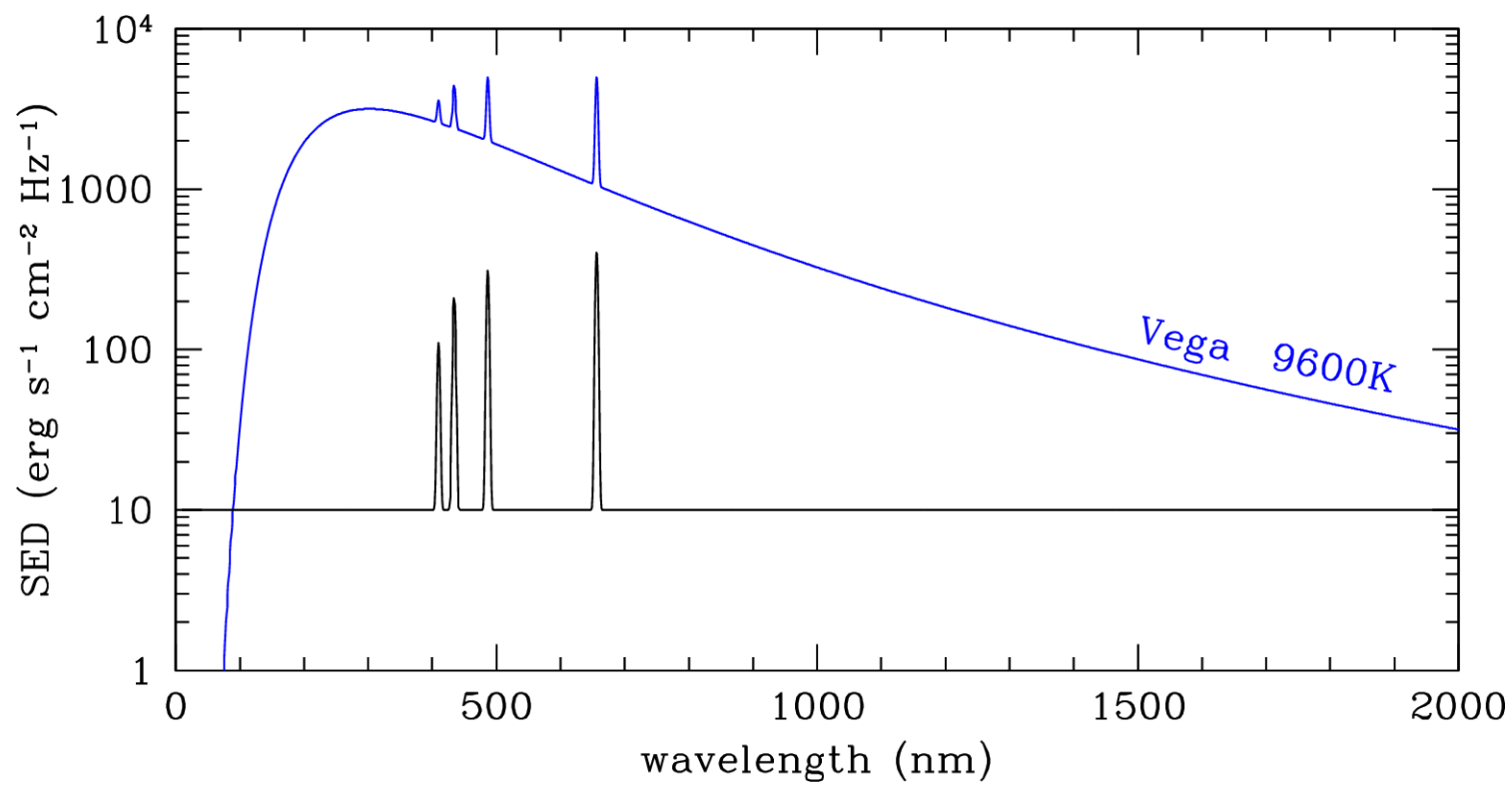
Linie absorpcyjne



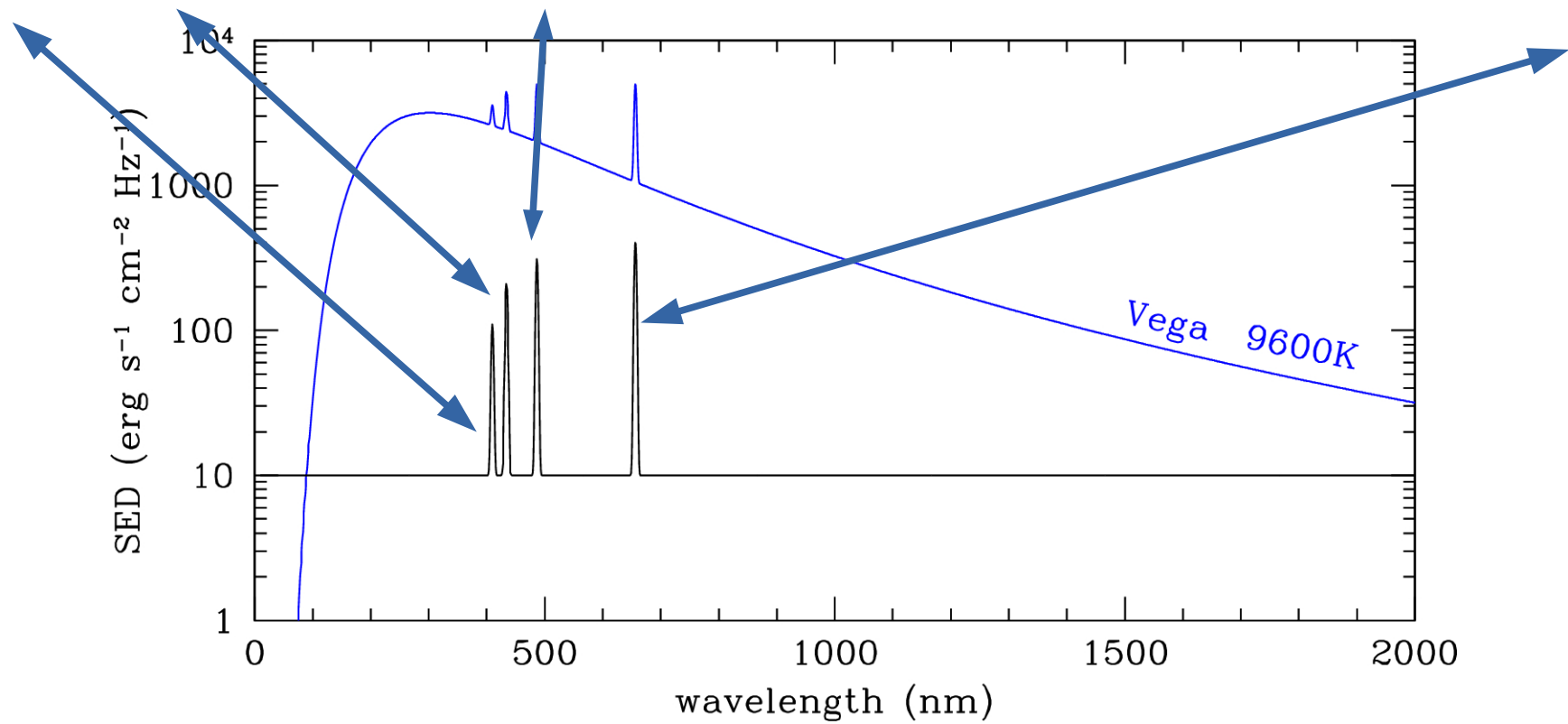
Widmo gwiazdy



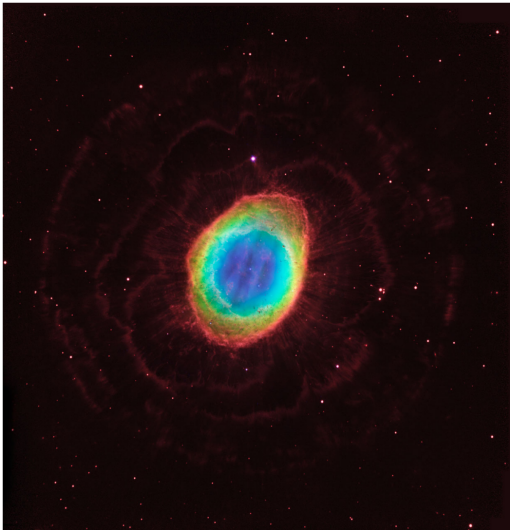
Linie emisyjne



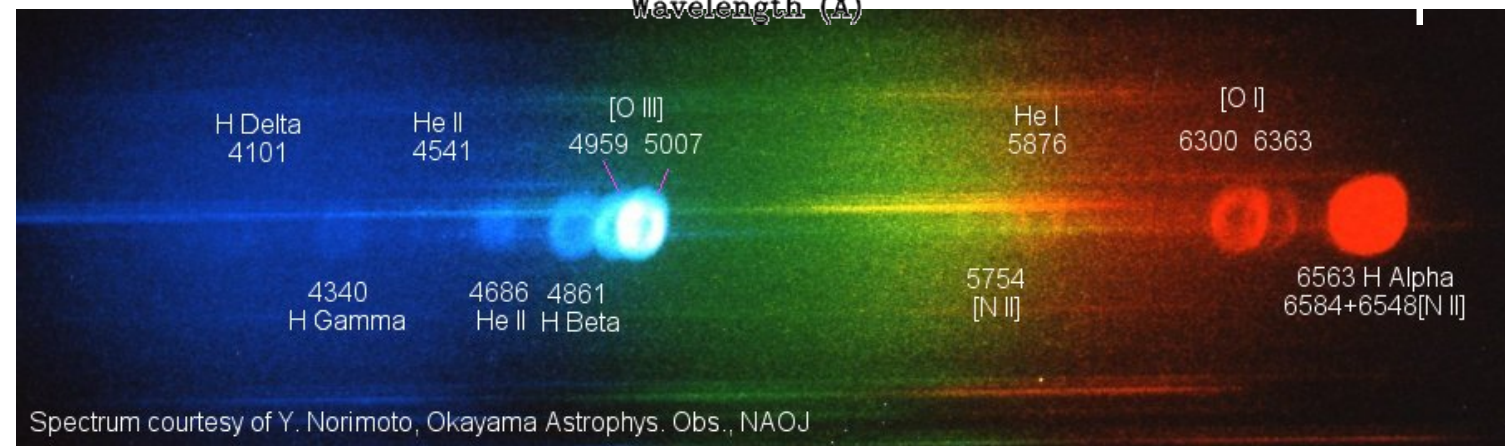
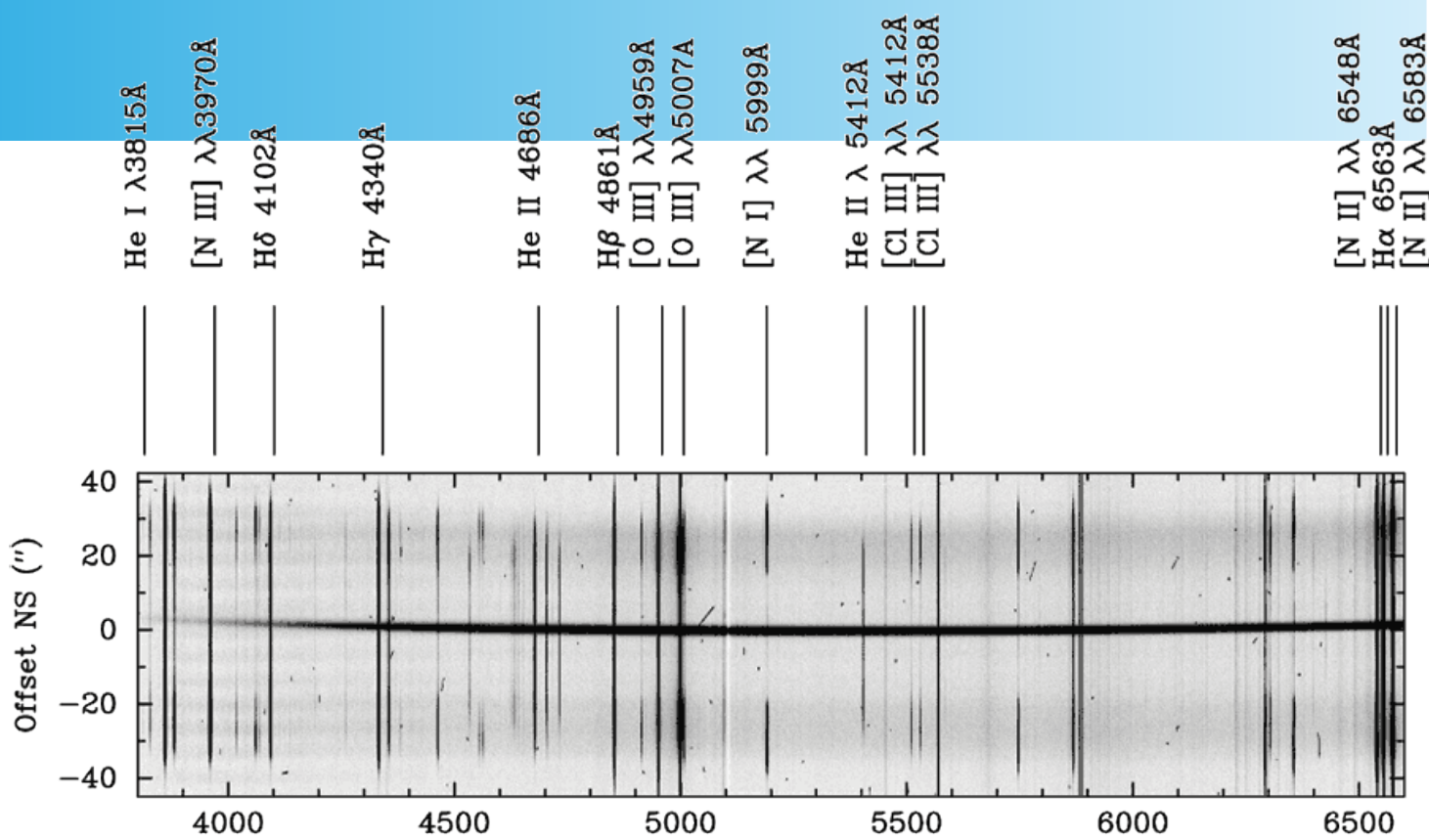
Linie emisyjne



M57 Spectrum - Blue Grism



M57

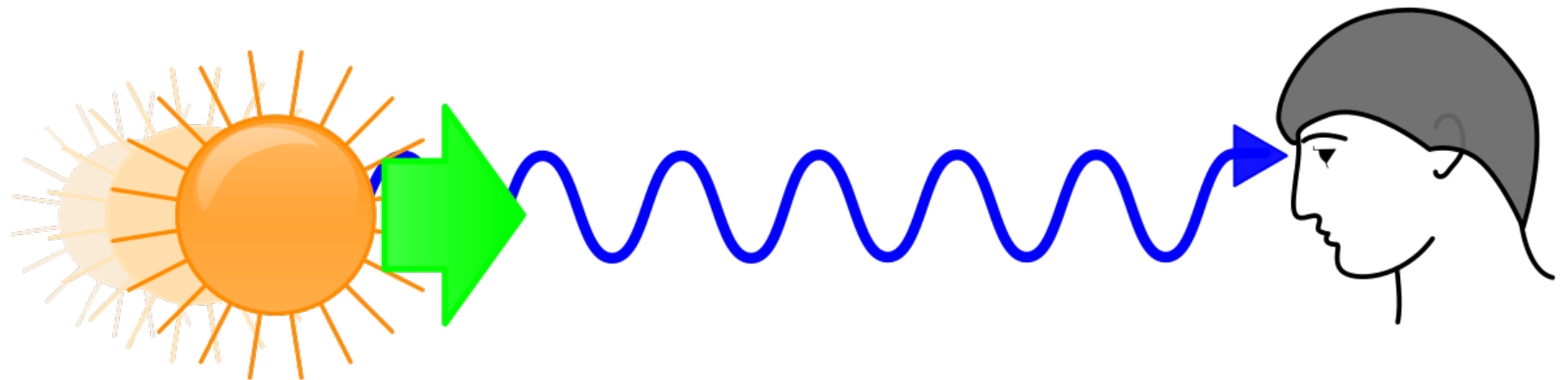
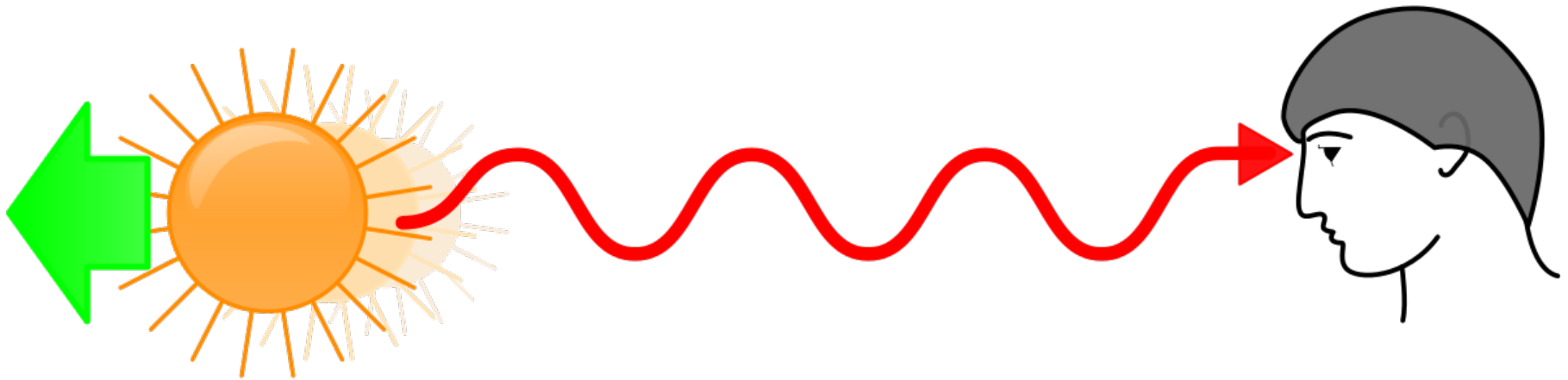


Spectrum courtesy of Y. Norimoto, Okayama Astrophys. Obs., NAOJ

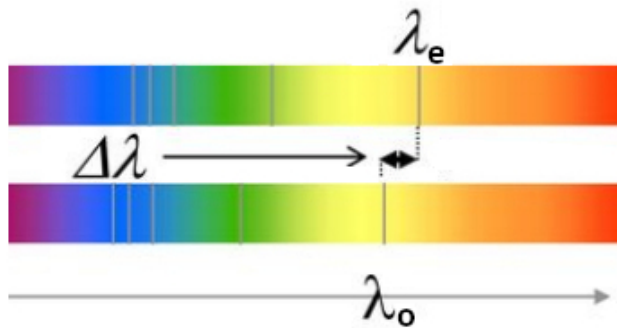
Linie wzbronione

- Oznaczamy je kwadratowymi nawiasami np. [OIII]
- Powstają w bardzo rzadkich gazach
- Czas pomiędzy zderzeniami atomów dużo dłuższy niż czas spontanicznej emisji fotonu

Efekt Dopplera



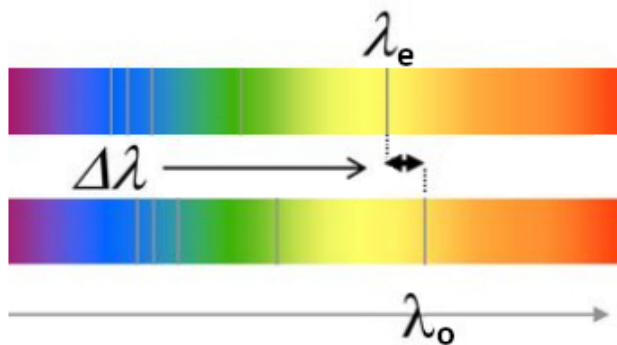
Blueshift / Redshift



$$\Delta\lambda = \lambda_o - \lambda_e$$

$\Delta\lambda$ is positive for redshift

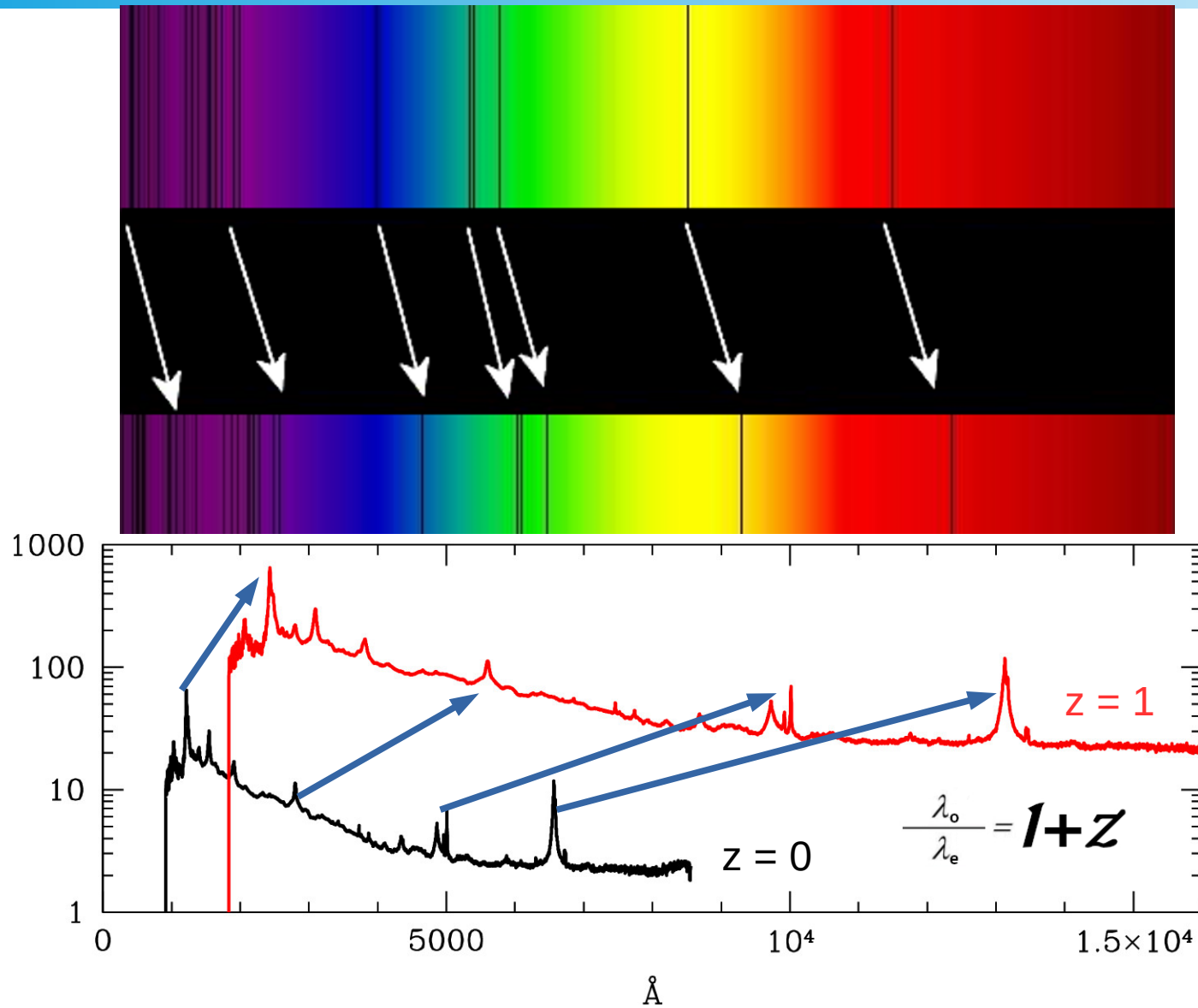
$\Delta\lambda$ is negative for blueshift



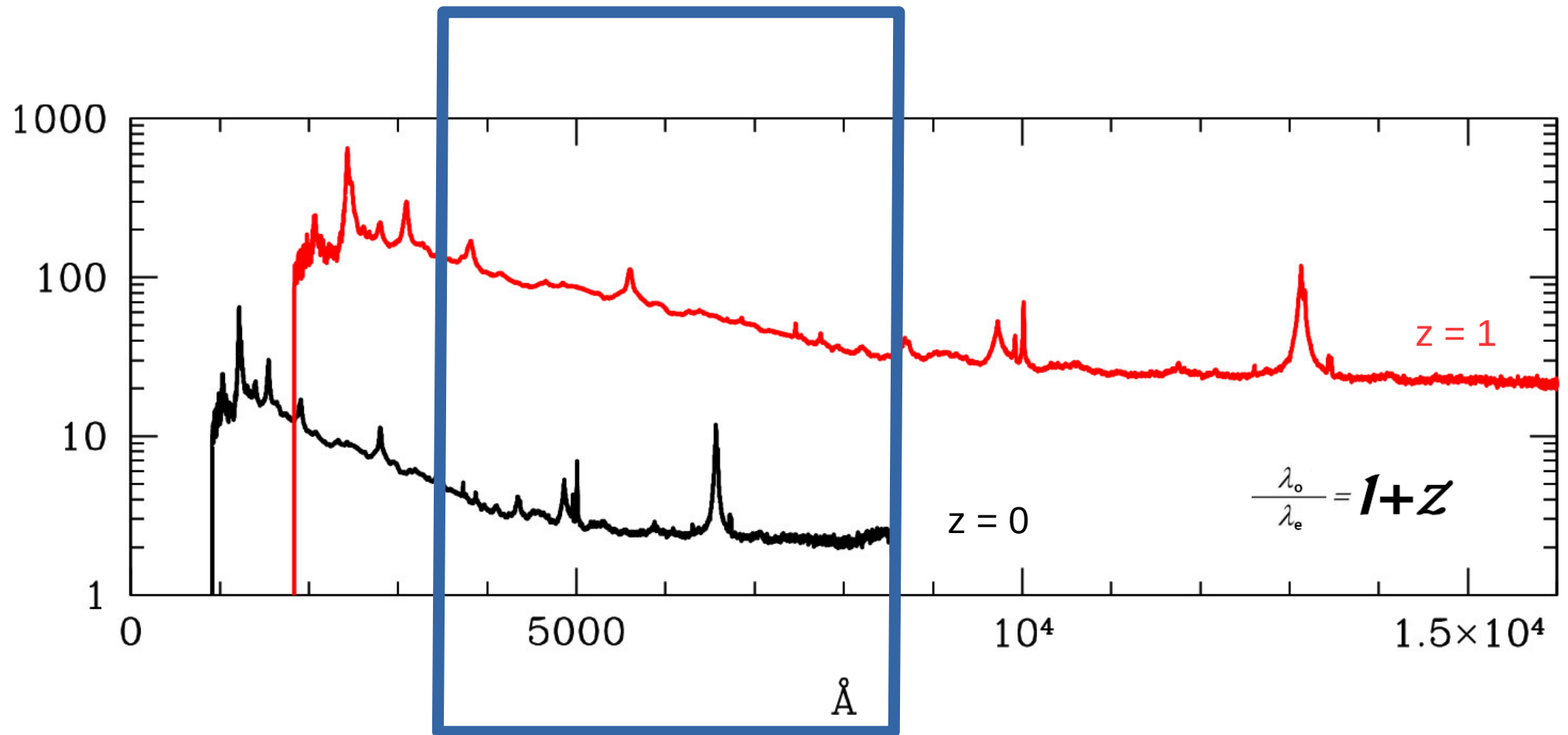
$$\frac{\lambda_o - \lambda_e}{\lambda_e} = z$$

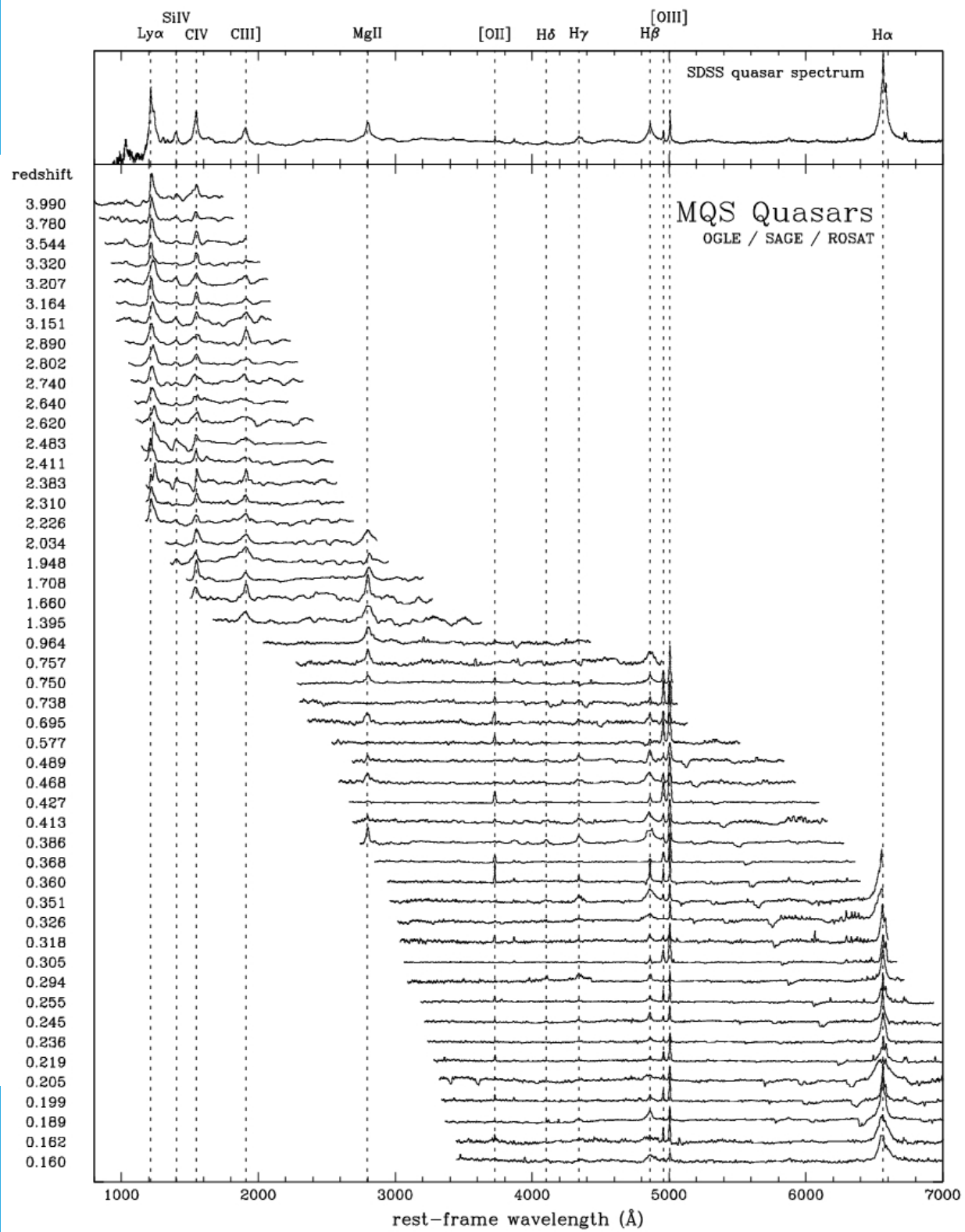
$$\frac{\lambda_o}{\lambda_e} = 1 + z$$

Redshift

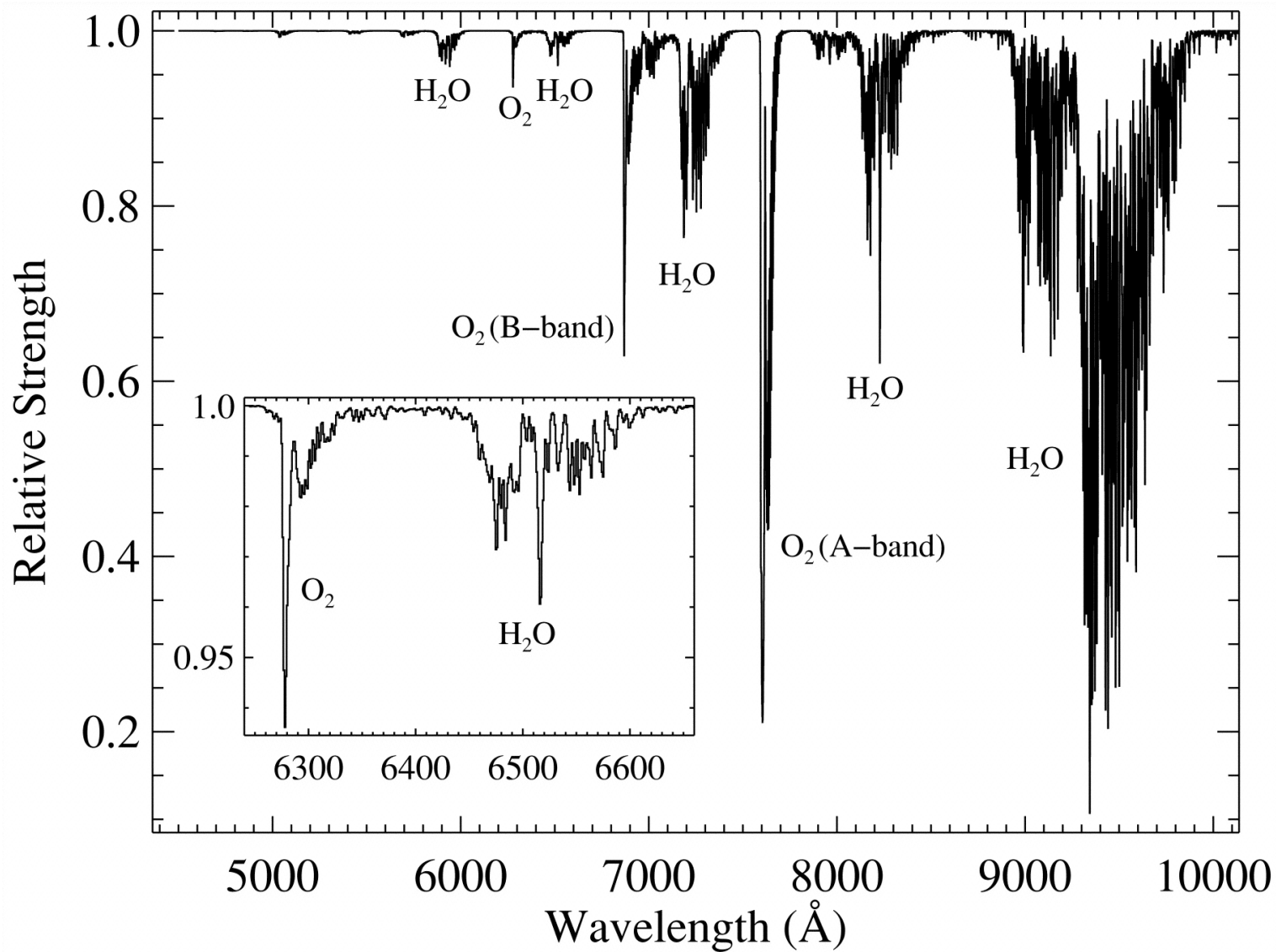


Zakres widmowy

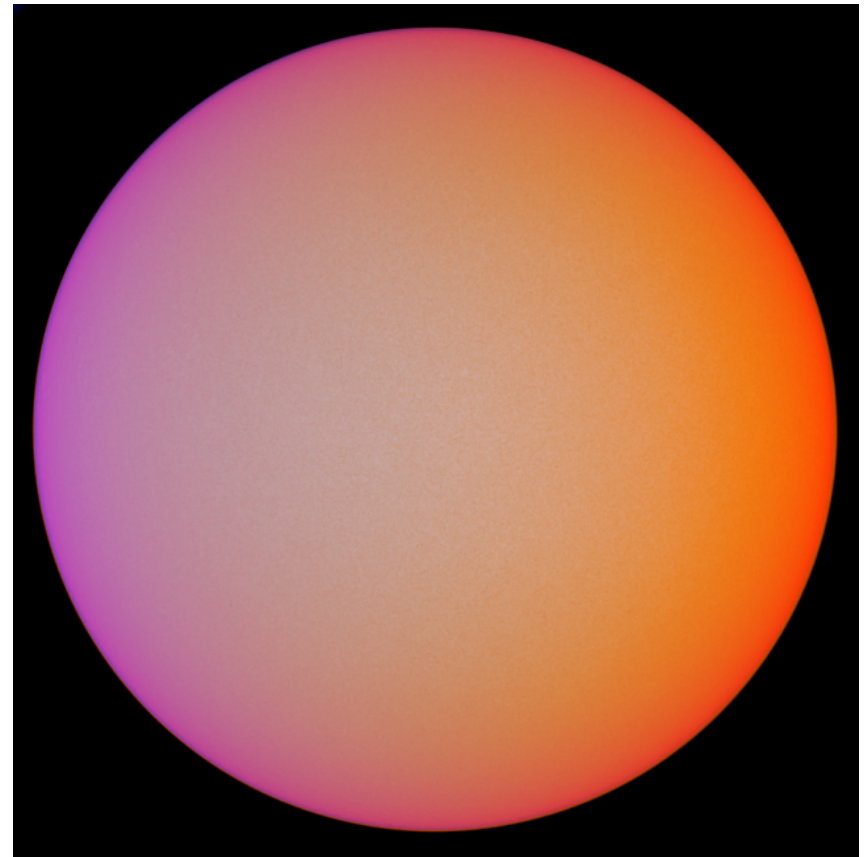
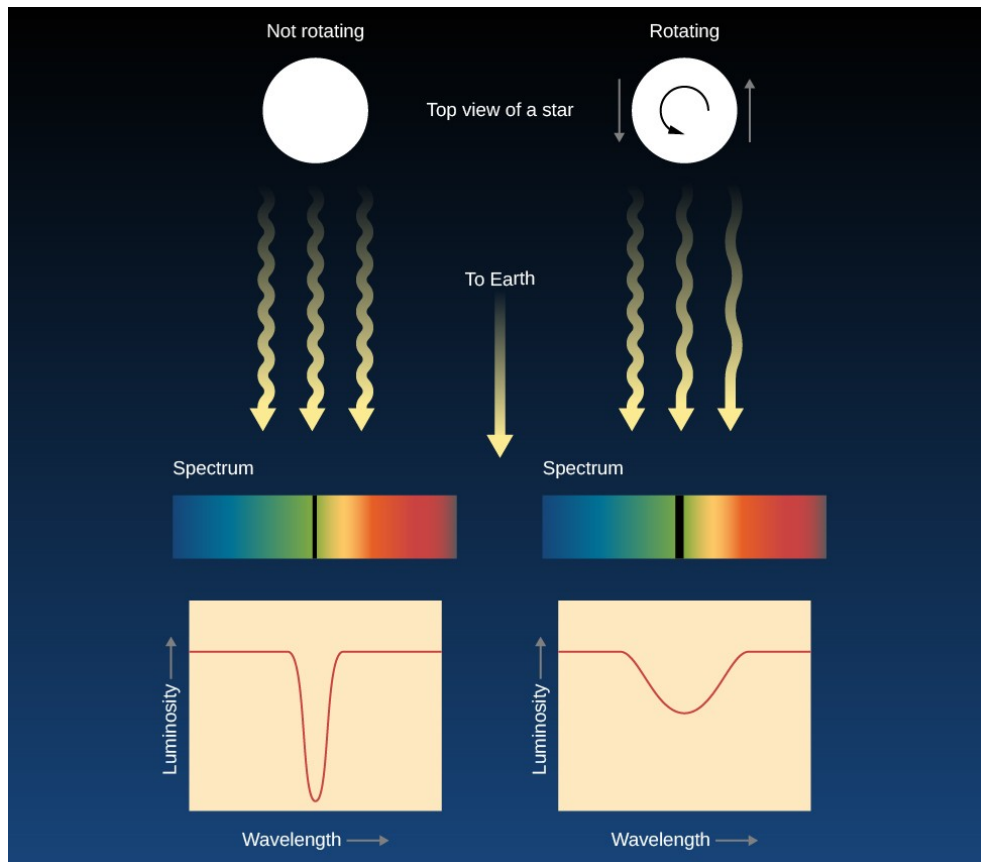




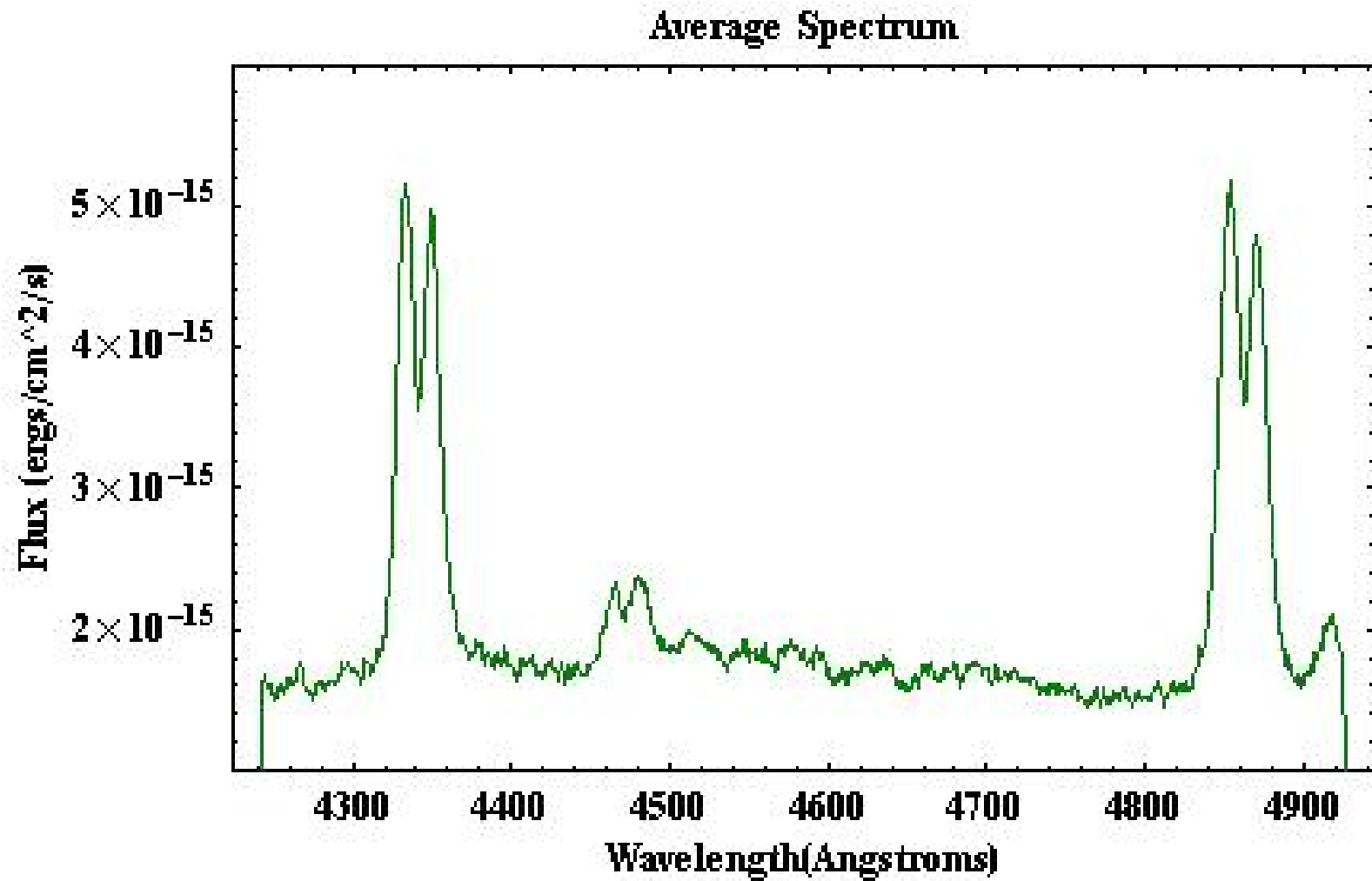
Linie Telluryczne



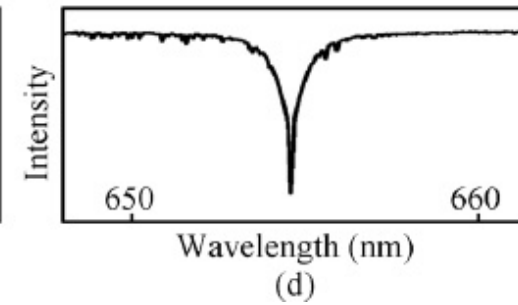
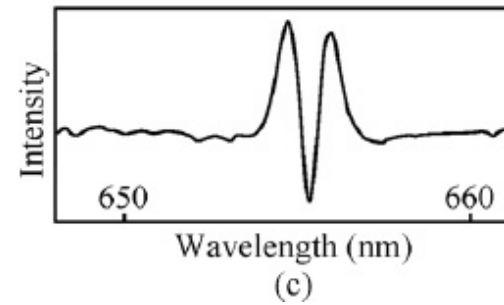
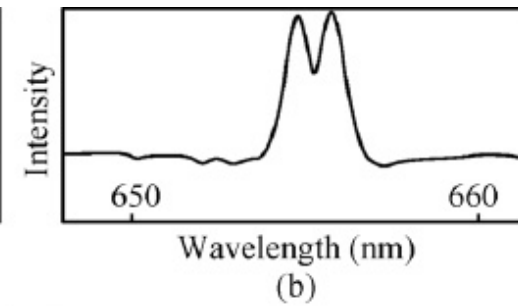
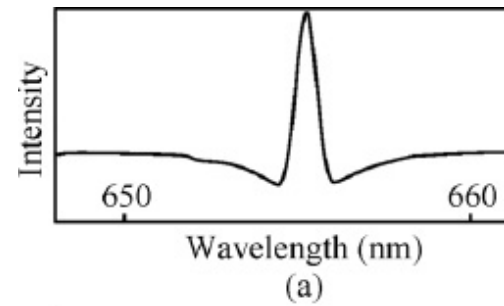
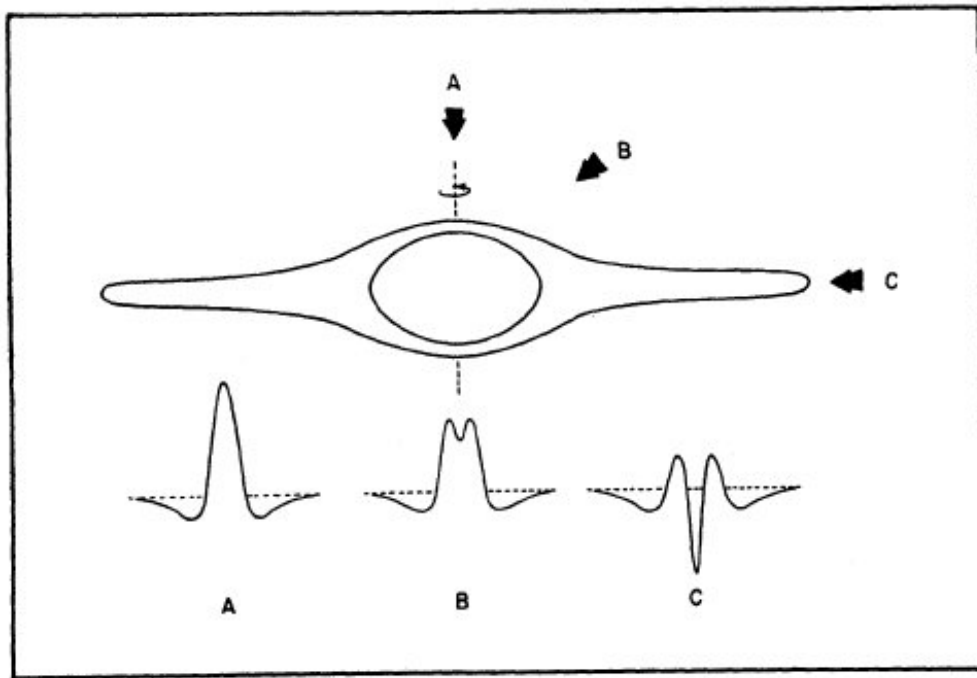
Poszerzenie rotacyjne linii



Dysk w CV



Gwiazdy Be (absorpcja emisji)



Profil P-Cygni

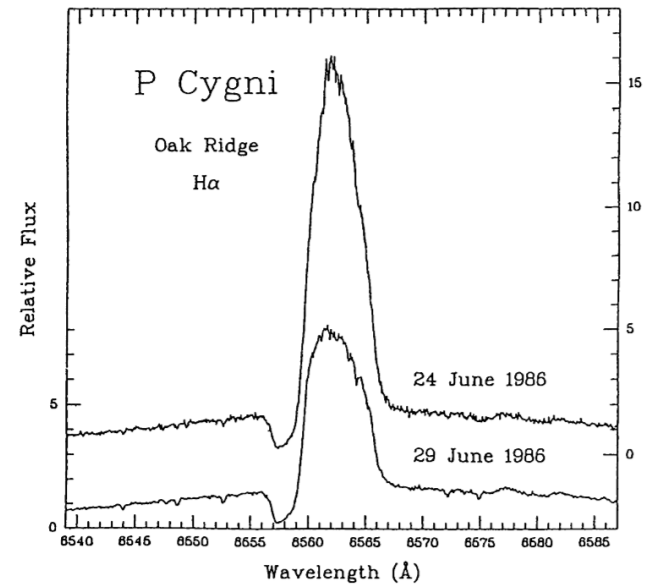
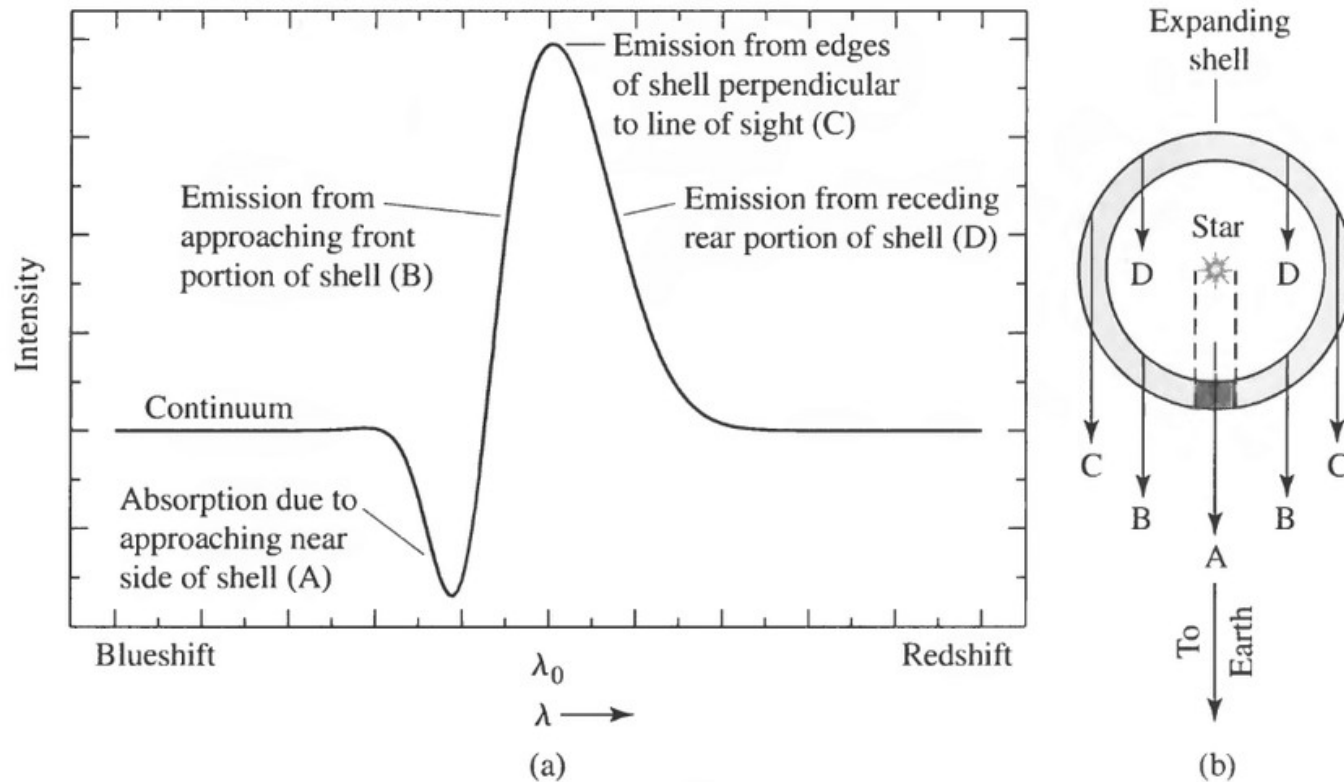
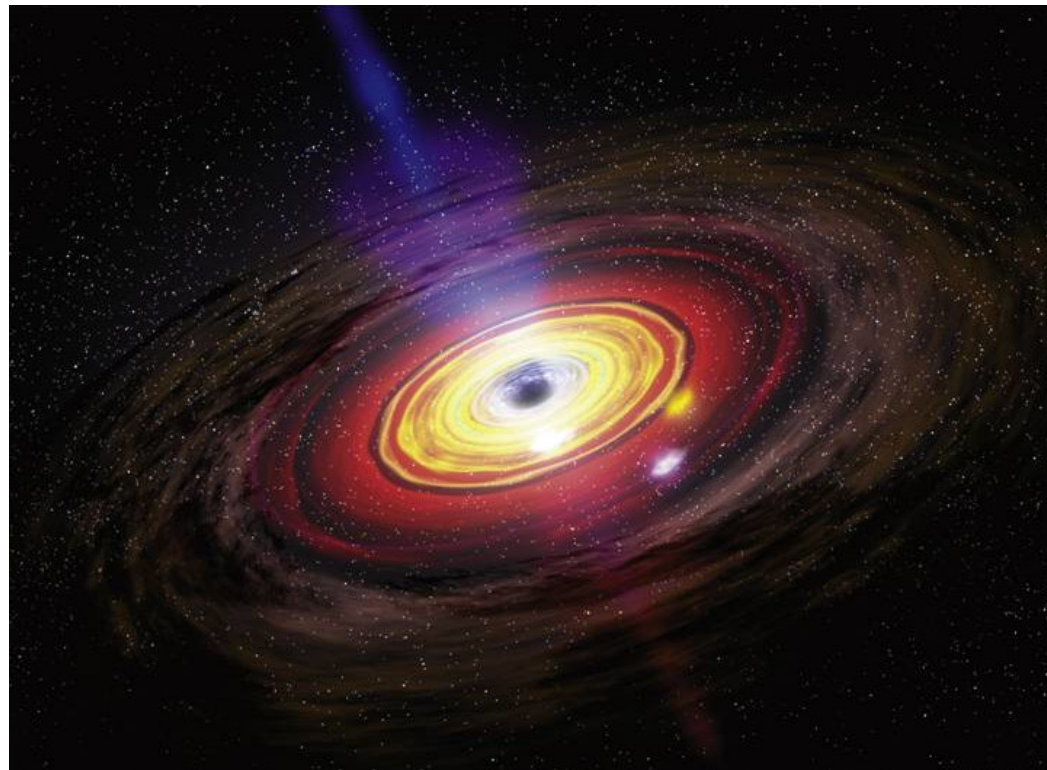
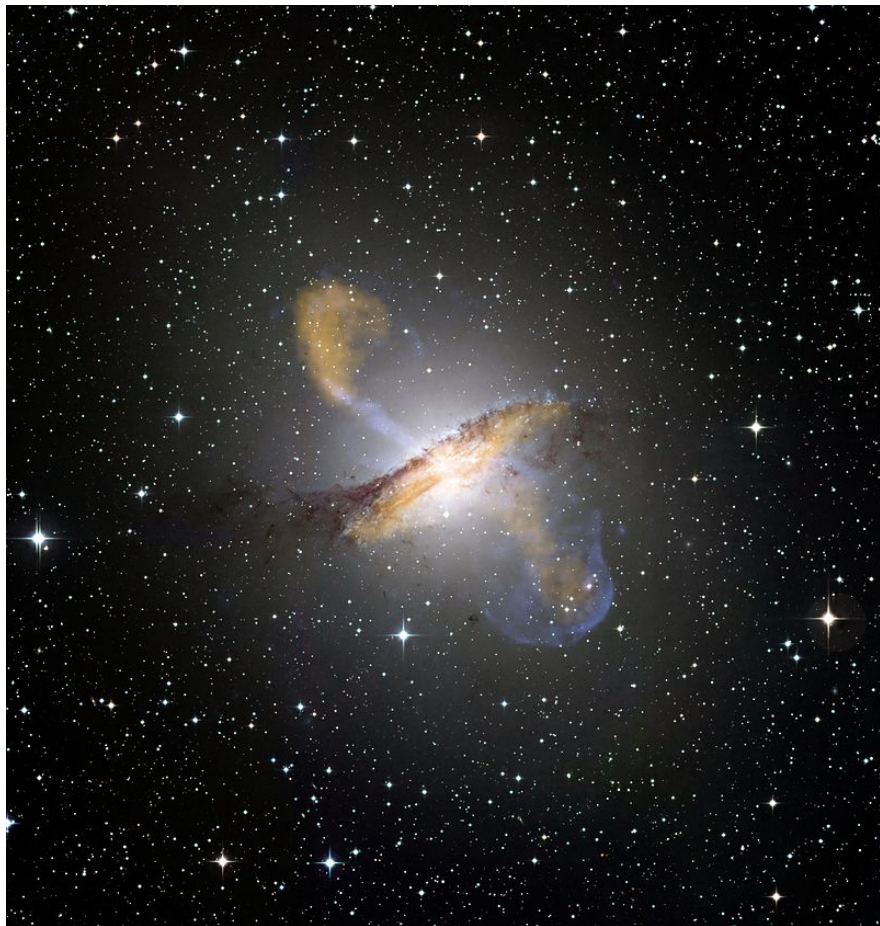
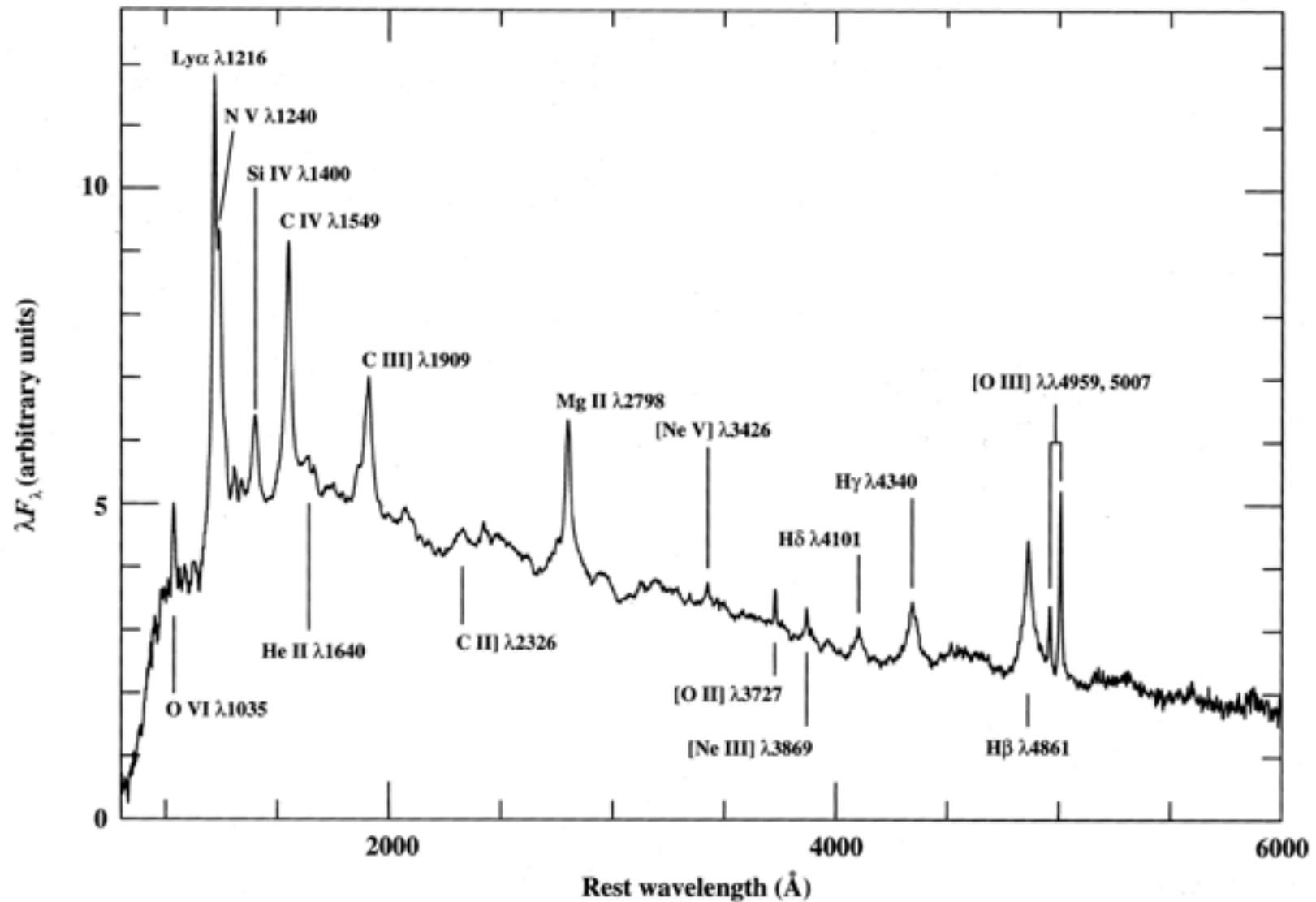


Figure 4. H-alpha profiles of P Cygni from Oak Ridge Observatory.

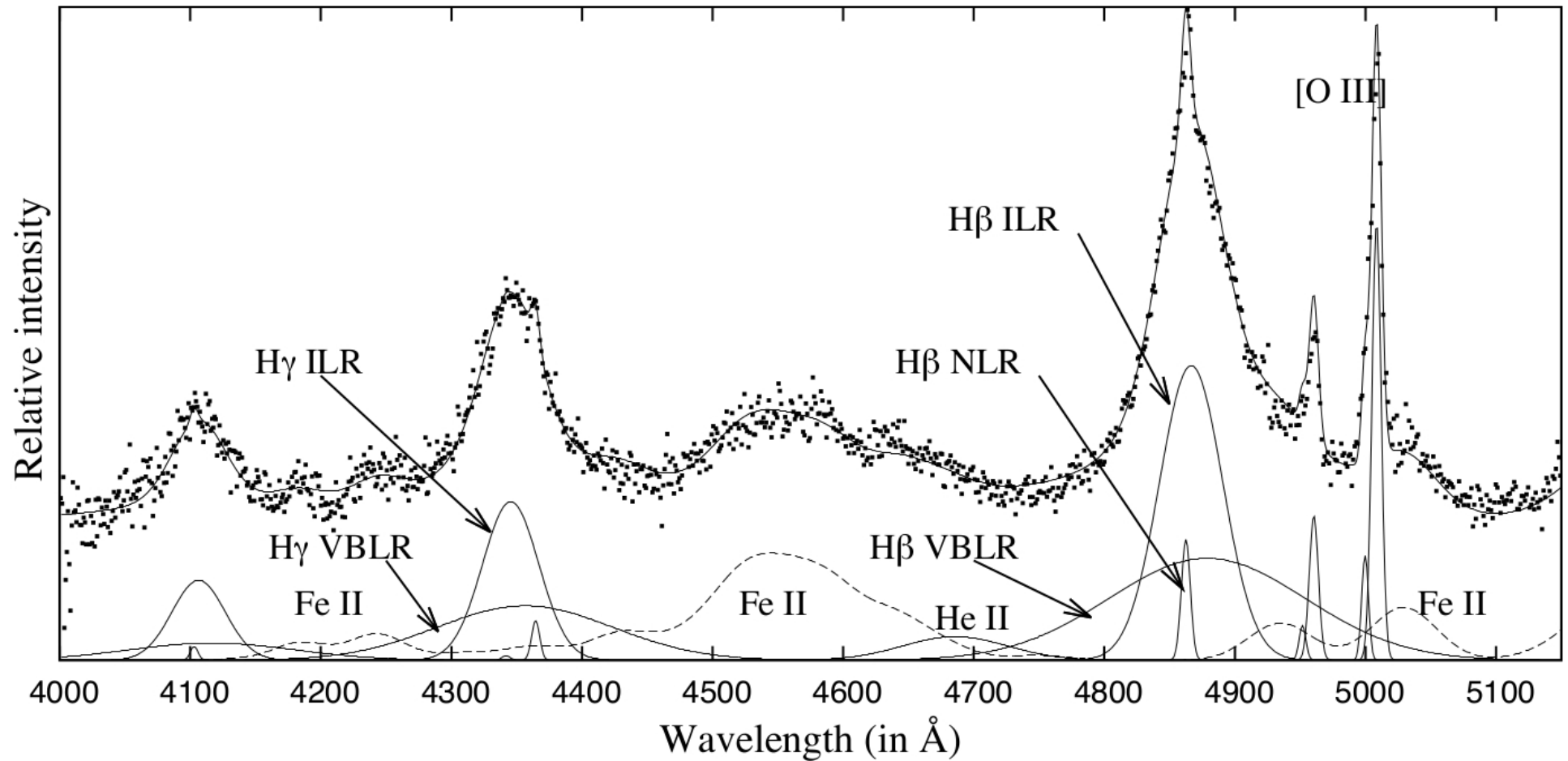
AGNy (poszerzenie linii)



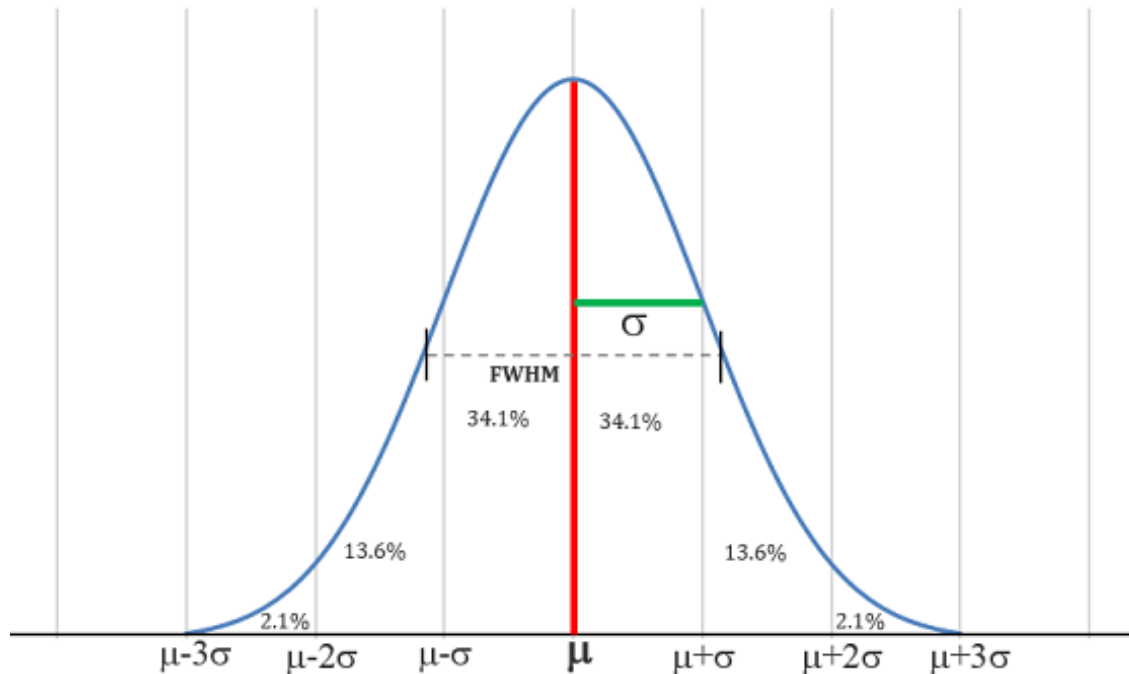
Widmo kwazara/AGNa



AGNy (poszerzenie linii)



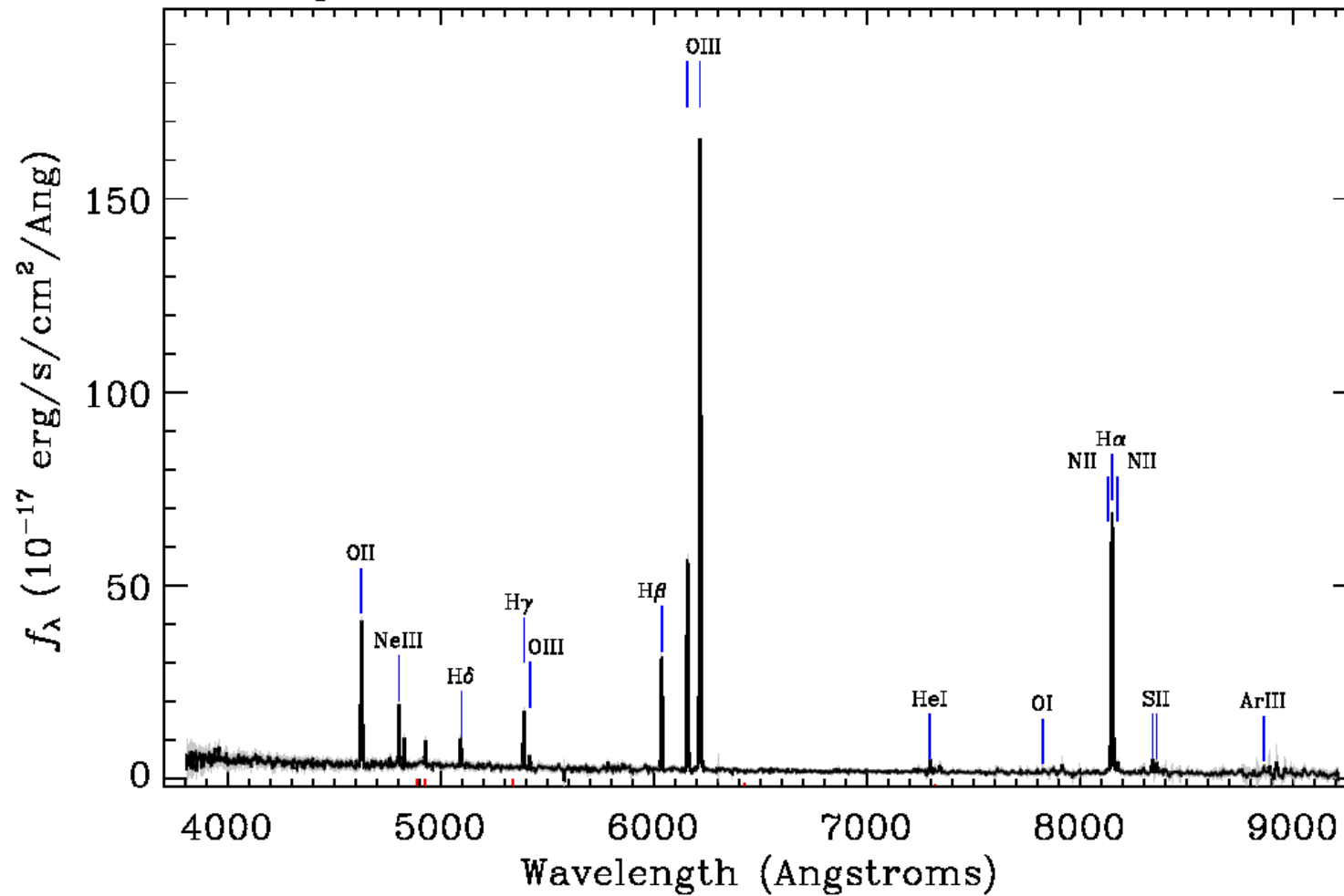
Pomiar masy czarnej dziury



$$\log\left(\frac{M_{\text{BH}}}{M_{\odot}}\right) = a + b \log\left(\frac{L \times 10^{-44}}{\text{erg s}^{-1}}\right) + 2 \log\left(\frac{\text{FWHM}}{\text{km s}^{-1}}\right)$$

Podstawowe linie

Survey: *sdss* Program: *legacy* Target: *SERENDIP_DISTANT_QSO_MAG_OUTLIER*
RA=173.26577, Dec=65.22815, Plate=597, Fiber=460, MJD=52059
 $z=0.24141 \pm 0.00001$ Class=GALAXY STARBURST
No warnings.



Podstawowe linie

- [SII] – 6716Å i 6732Å
- H α – 6563Å
- [NII] – 6548 i 6584Å
- [OIII] – 4959 i 5007Å
- H β – 4861Å
- [OII] – 3727Å
- MgII – 2800Å
- Ly α – 1216Å

Metalliczność

For the line fluxes, we will be using the following notations throughout the paper: $R_2 = I_{[\text{O II}]\lambda 3727 + \lambda 3729} / I_{\text{H}\beta}$, $R_3 = I_{[\text{O III}]\lambda 4959 + \lambda 5007} / I_{\text{H}\beta}$, $R = I_{[\text{O III}]\lambda 4363} / I_{\text{H}\beta}$, and $R_{23} = R_2 + R_3$. With these definitions, the excitation parameter P can then be expressed as $P = R_3 / (R_2 + R_3)$.

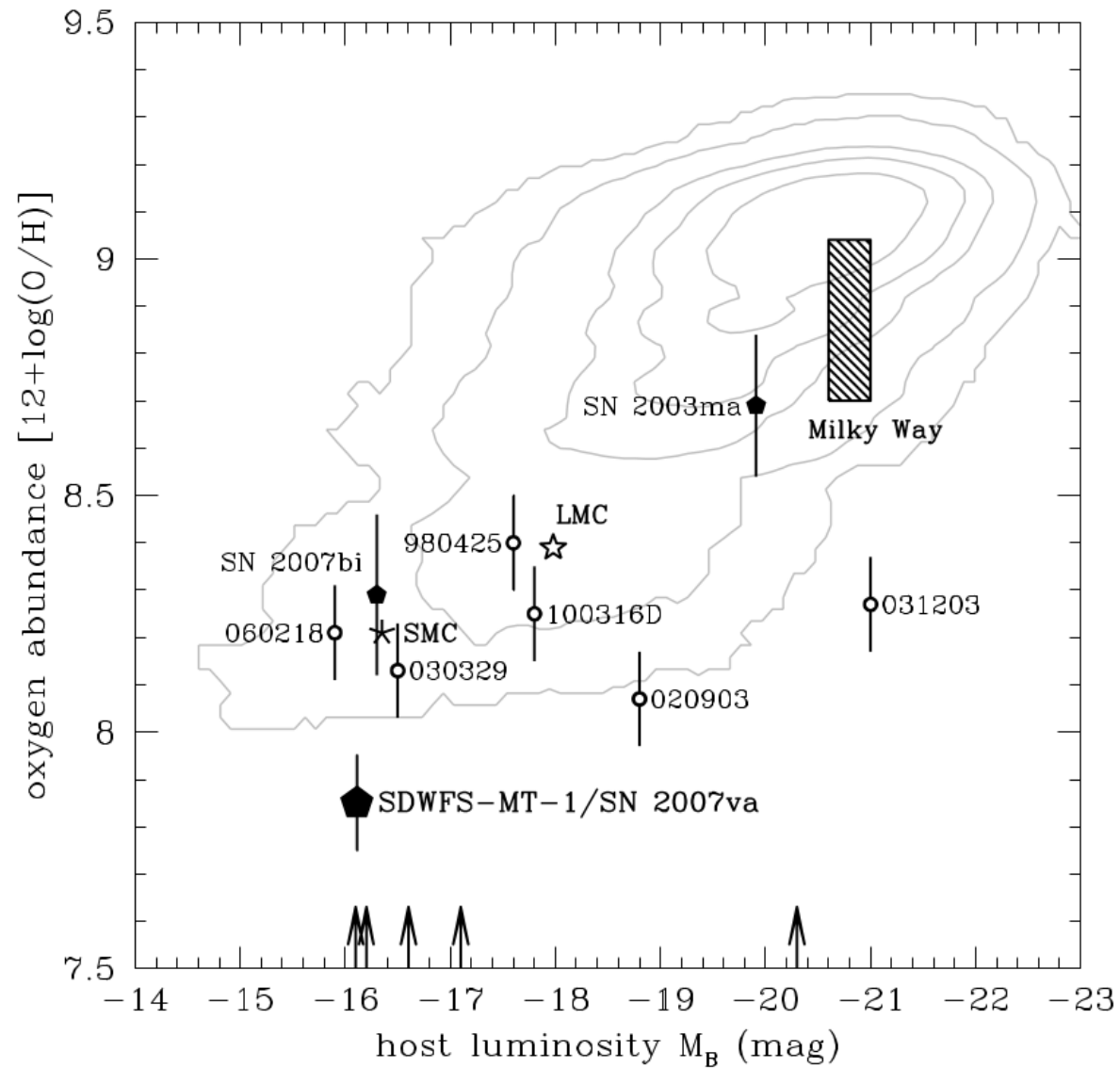
$$\begin{aligned} \text{Z94: } \log(\text{O}/\text{H}) + 12 = & 9.265 - 0.33 R_{23} - 0.202 R_{23}^2 \\ & - 0.207 R_{23}^3 - 0.333 R_{23}^4, \quad (8) \end{aligned}$$

$$\begin{aligned} \text{M91(u): } \log(\text{O}/\text{H}) + 12 = & 12.0 - 4.944 + 0.767 R_{23} \\ & + 0.602 R_{23}^2 \\ & - y(0.29 + 0.332 R_{23} \\ & - 0.331 R_{23}^2), \quad (9) \end{aligned}$$

where

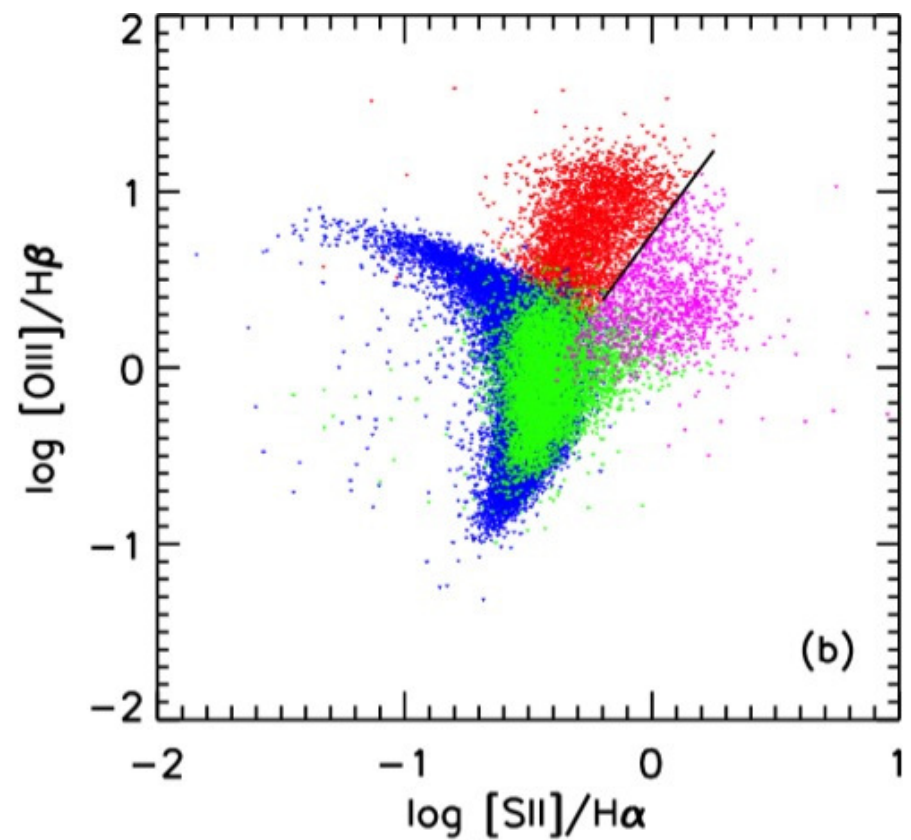
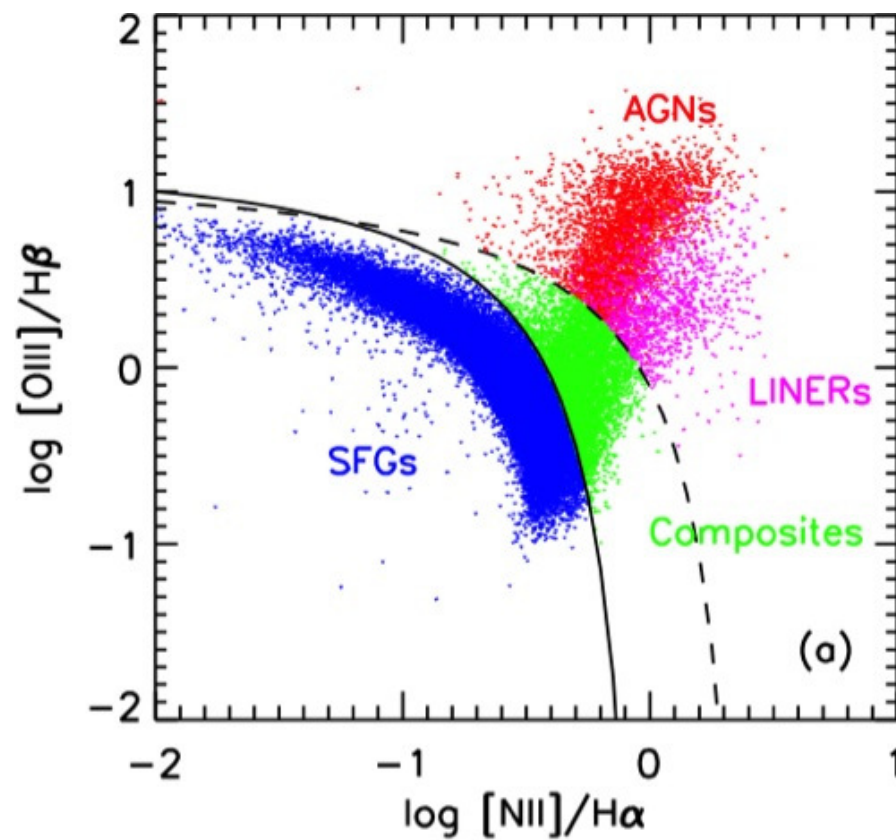
$$y = \log\left(\frac{[\text{O III}] \lambda\lambda 4959, 5007}{[\text{O II}] \lambda 3727}\right), \quad (10)$$

Metalliczność

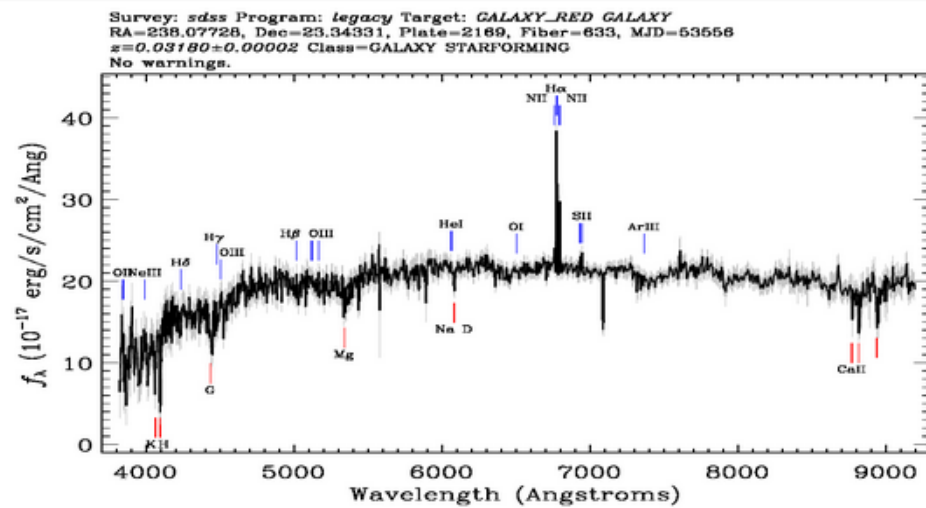
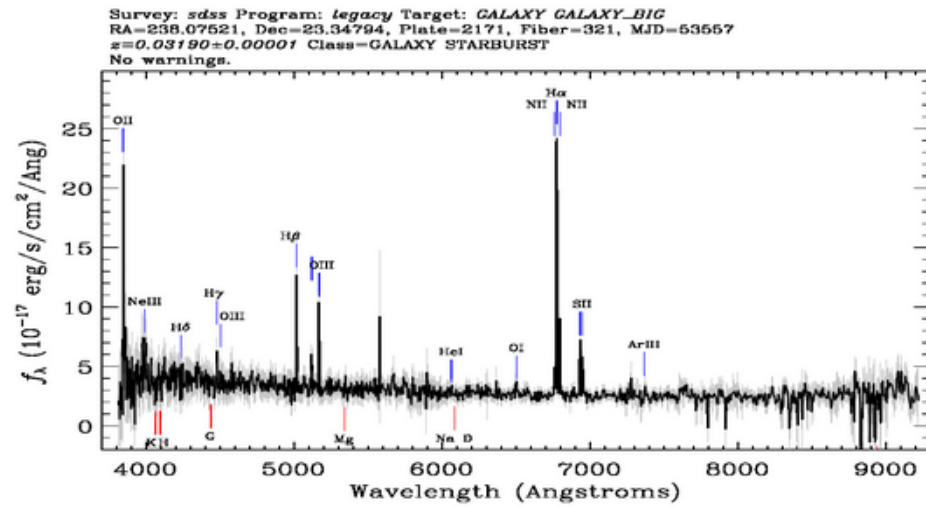
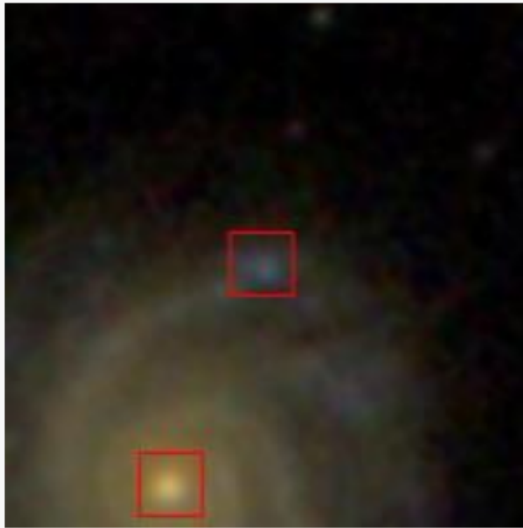


BPT diagram

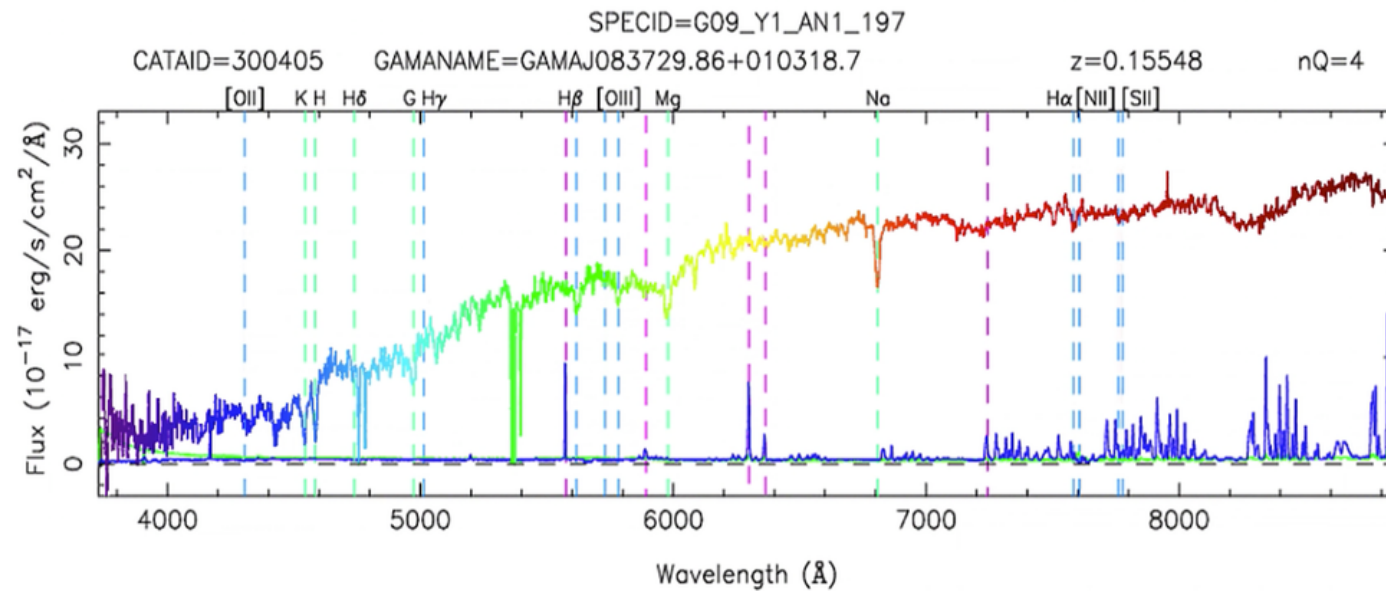
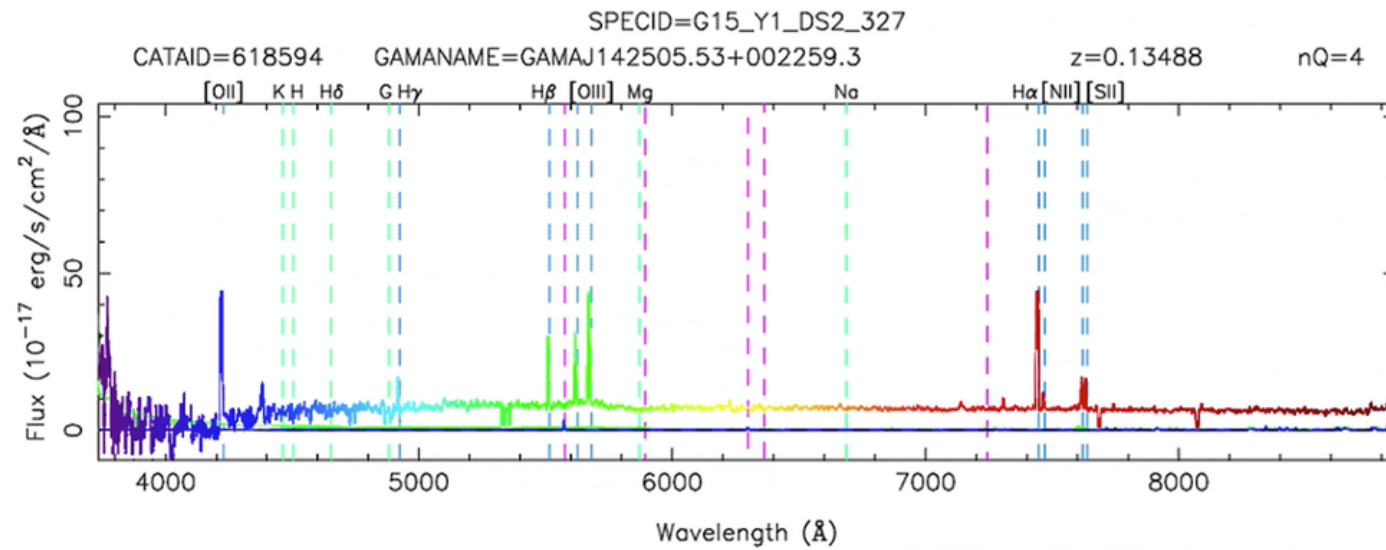
BPT = Baldwin, Phillips & Terlevich (1981)



Widmo galaktyki



Widma galaktyk



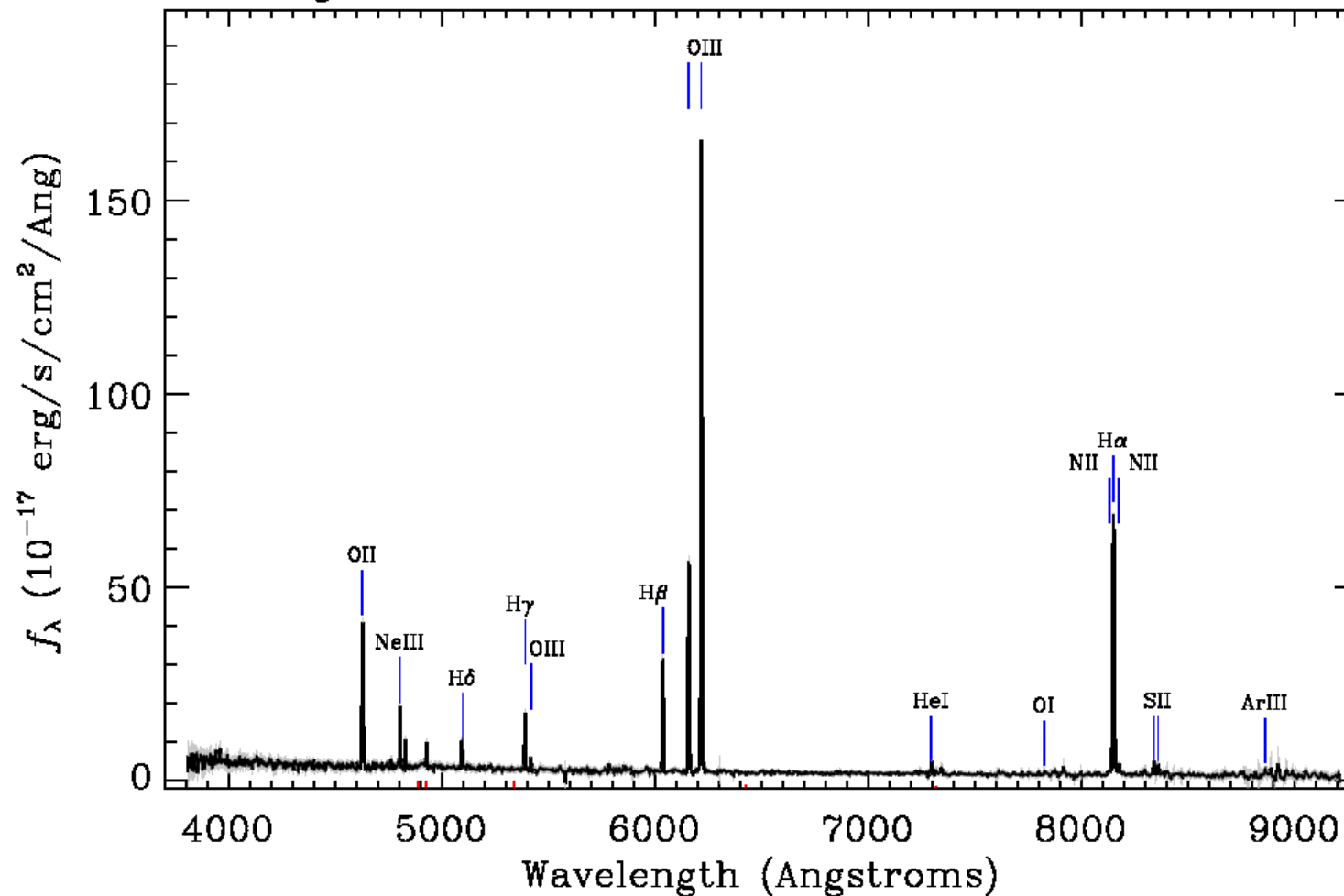
Galaktyka formująca gwiazdy

Survey: *sdss* Program: *legacy* Target: *SERENDIP_DISTANT_QSO_MAG_OUTLIER*

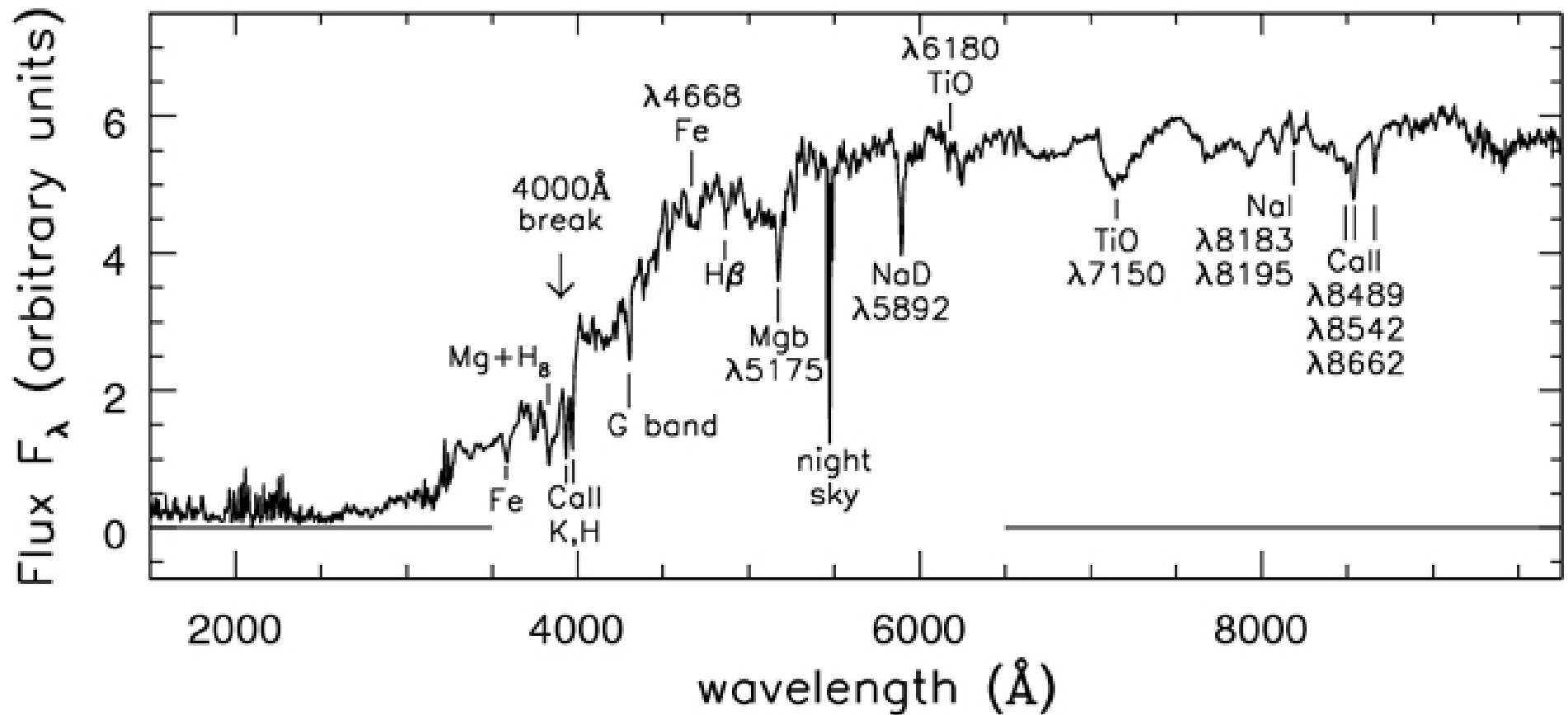
RA=173.26577, Dec=65.22815, Plate=597, Fiber=460, MJD=52059

$z=0.24141 \pm 0.00001$ Class=GALAXY STARBURST

No warnings.



Galaktyka eliptyczna



Widmo kwazara/AGNa

